

Publication Year	2017
Acceptance in OA	2024-02-16T16:00:27Z
Title	Geant4 simulations of soft proton scattering in X-ray optics. A tentative validation using laboratory measurements
Authors	FIORETTI, VALENTINA, MINEO, TERESA, BULGARELLI, ANDREA, Dondero, Paolo, Ivanchenko, Vladimir, LEI, FAN, LOTTI, Simone, MACCULI, CLAUDIO, Mantero, Alfonso
Publisher's version (DOI)	10.1007/s10686-017-9559-9
Handle	http://hdl.handle.net/20.500.12386/34766
Journal	EXPERIMENTAL ASTRONOMY
Volume	44

# Geant4 simulations of soft proton scattering in X-ray optics

A tentative validation using laboratory measurements

Valentina Fioretti · Teresa Mineo · Andrea Bulgarelli · Paolo Dondero · Vladimir Ivanchenko · Fan Lei · Simone Lotti · Claudio Macculi · Alfonso Mantero on behalf of the AREMBES collaboration

Received: date / Accepted: date

Abstract Low energy protons (< 300 keV) can enter the field of view of X-ray space telescopes, scatter at small incident angles, and deposit energy on the detector. This phenomenon can cause intense background flares at the focal plane decreasing the mission observing time (e.g. the XMM-Newton mission) or in the most extreme cases, damaging the X-ray detector. A correct modelization of the physics process responsible for the grazing angle scattering processes is mandatory to evaluate the impact of such events on the performance (e.g. observation time, sensitivity) of future X-ray telescopes as the ESA ATHENA mission. The Remizovich model describes particles reflected by solids at glancing angles in terms of the Boltzmann transport equation using the diffuse approximation and the model of continuous slowing down in energy. For the first time this solution, in

V. Fioretti

 ${\it INAF/IASF}$ Bologna, Via P. Gobetti 101, 40129, Bologna, Italy E-mail: fioretti@iasfbo.inaf.it

T. Mineo

INAF/IASF Palermo, Via Ugo La Malfa 153, 90146 Palermo, Italy

A. Bulgarelli

INAF/IASF Bologna, Via P. Gobetti 101, 40129, Bologna, Italy

P. Dondero

SWHARD srl, Via Greto di Cornigliano, 6, 16152, Genova, Italy

V. Ivanchenko

CERN, Geneve 23, 1211, Switzerland

Geant<br/>4 Associate International Ltd, 9 Royd Terrace, Hebden Bridge HX7 7BT, United Kingdom

F. Lei

RadMod Research, 25 Parkway, Camberley, Surrey, GU15 2PD, United Kingdom

S. Lotti

INAF/IAPS, Via del Fosso del Cavaliere, 100, 00133 Roma, Italy

C. Macculi

INAF/IAPS, Via del Fosso del Cavaliere, 100, 00133 Roma, Italy

A. Mantero

SWHARD srl, Via Greto di Cornigliano, 6, 16152, Genova, Italy

the approximation of no energy losses, is implemented, verified, and qualitatively validated on top of the Geant4 release 10.2, with the possibility to add a constant energy loss to each interaction. This implementation is verified by comparing the simulated proton distribution to both the theoretical probability distribution and with independent ray-tracing simulations. Both the new scattering physics and the Coulomb scattering already built in the official Geant4 distribution are used to reproduce the latest experimental results on grazing angle proton scattering. At 250 keV multiple scattering delivers large proton angles and it is not consistent with the observation. Among the tested models, the single scattering seems to better reproduce the scattering efficiency at the three energies but energy loss obtained at small scattering angles is significantly lower than the experimental values. In general, the energy losses obtained in the experiment are higher than what obtained by the simulation. The experimental data are not completely representative of the soft proton scattering experienced by current X-ray telescopes because of the lack of measurements at low energies (< 200 keV) and small reflection angles, so we are not able to address any of the tested models as the one that can certainly reproduce the scattering behavior of low energy protons expected for the ATHENA mission. We can, however, discard multiple scattering as the model able to reproduce soft proton funnelling, and affirm that Coulomb single scattering can represent, until further measurements at lower energies, the best approximation of the proton scattered angular distribution at the exit of X-ray optics.

Keywords Geant $4 \cdot \text{soft protons} \cdot X$ -ray telescopes  $\cdot ATHENA$ 

#### 1 Introduction

Charged particles can pose a significant radiation threat to the on-board electronic systems of X-ray space missions, depending on the telescope orbit. This phenomenon is particularly complex for grazing incident X-ray telescopes as the NASA Chandra X-ray Observatory (CXO) [1] and the ESA XMM-Newton [2] telescopes, launched in July and September 1999 respectively, and currently operating in a highly eccentric orbit that crosses the radiation belt in the  $(5-10) \times 10^3$  km altitude range. Both telescopes carry Wolter-I type mirrors to focus X photons through grazing angle reflection to the detection plane. The capability of X-ray optics to focus electrons was already known before the launch of the two missions. For this reason, X-ray telescopes were equipped with magnetic diverters that deflect the electron paths outside the detection plane (see e.g. [3]). However, the loss of charge transfer efficiency suffered by the Chandra Advanced CCD Imaging Spectrometer (ACIS [4]) Front-Illuminated (FI) CCDs after the first radiation belt passages revealed that low energy protons can also be reflected by the X-ray mirror shells and reach the focal plane. With energies between 1-300 keV, these so-called "soft protons" caused serious damages to the FI CCD at the focal plane of CXO and the overall mission performance of XMM-Newton. In fact, XMM-Newton's filter wheel completely blocks the EPIC [5] field of view when crossing the radiation belts protecting the detectors from damage. Unfortunately above the radiation belt limit, where the instruments are fully operative, the soft proton funnelling is still observed by the XMM-Newton detectors in the form of sudden

increases in the background level. Such soft proton flaring events can prevail over the quiescent background level up to orders of magnitude, affecting 30-40% of XMM-Newton observing time [6]. Soft proton flares are extremely unpredictable in duration, lasting from  $\sim 100~\rm s$  to hours [7], and generate an average count rate, in all three CCDs, of 2-2.5 prot. cm $^{-2}~\rm s^{-1}$  [8]. Several studies (see e.g. 6) have proven the solar origin of this damaging background. Missions operating in low Earth orbit (e.g. Swift [9], Suzaku [10]) do not suffer from soft proton flares, thanks to the geomagnetic shield. Soft protons populating the outer magnetosphere, including the magnetotail, the magnetosheath and the solar wind, both in the form of a steady flux and violent Coronal Mass Ejections, can instead increase the X-ray residual background level and even threaten the observation itself [11].

Future X-ray focusing telescopes operating outside the radiation belts (e.g. the ESA Athena [12] mission and the eROSITA [13] instrument on-board the Russian Spektr-RG observatory, both to be placed in L2 orbit) will also be affected by soft proton contaminations.

A validated physics model to describe the angular and energy proton distribution at the exit of the optics is mandatory for a correct evaluation of the impact of soft proton events to the mission performance. Despite many solutions proposed so far to explain the physics interaction behind the soft proton grazing angle scattering (see e.g. [14]), the difficulty in implementing dedicated models prevented Monte Carlo codes from developing physically sound models. We implement in the Geant4 release 10.2 both the Firsov distribution following the work of [15] and for the first time the bivariate distribution described by Remizovich in its elastic approximation (see Sec. 5). After a dedicated verification of the implementation performed comparing the results with the analytical model and the ray-tracing simulation, we use the latest scattering measurements obtained at 250, 500, and 1000 keV proton energy by [16] on eRosita shell samples to accurately compare the measured energy, angular, and intensity distribution of protons with the ones predicted from both the new models and the Geant4 default library. The physics validation of a soft proton scattering model is achieved if we are able to reproduce, within an acceptable uncertainty level, the experimental data.

## 2 Experimental data

The experiment of [16] evaluates the scattering efficiency, in  ${\rm sr}^{-1}$ , and the energy loss of protons at 250, 500, and 1000 keV interacting with eRosita shell samples at glancing angles in the  $0.3^{\circ}-1.2^{\circ}$  range. The general set-up consists of a proton beam line produced by a ion accelerator facility hitting the X-ray mirror shell sample at different angles and then collected by a shiftable proton detector:

- the incident angle has a precision, in terms of tilt angle, of 0.006°;
- the eRosita target shell is composed by a Nickel substrate of 270  $\mu m$  coated by 50 nm of Gold and different sample sizes, ranging from a length of 10 cm to 12 cm, were used with no impact on the measurements;
- the proton detector consists of a 8 mm wide Silicon surf barrier detector characterized by a detection efficiency of almost 100% and an overall accuracy within  $\pm 10~{\rm keV}.$

The 1.2 mm aperture at the proton detection point, placed at a distance of 933 mm from the target, defines a proton collection solid angle  $\Omega \simeq 1.3~\mu sr.$  The

scattering angle is defined in the reference as  $\theta + \theta_0$ , the sum of the scattering and incident polar angles. The proton scattering efficiency  $\eta(\phi,\theta)$ , in sr<sup>-1</sup>, is computed by dividing the number of detected protons  $N_{\rm det}$  by the number of incident protons in the target,  $N_{\rm inc}$ , and the aperture solid angle  $\Omega$ :

$$\eta(\phi, \theta) = \frac{N_{\text{det}}}{N_{\text{inc}} \times \Omega} .$$
(1)

The incident angles, the resulting scattering efficiency values and the energy losses are taken as reference values in the validation test of Sec. 6.

#### 3 Investigated theoretical models

#### 3.1 Remizovich model

The analytical model of [17,18] describes particles reflected by solids at glancing angles in terms of the Boltzmann transport equation using the diffuse approximation and the model of continuous slowing down in energy. According to their model, the proton energy loss is peaked at about 10-20% of the initial energy  $E_0$ , i.e. ranges from 5-10 keV at 50 keV to 50-100 keV at 500 keV. The mathematical form of the Remizovich model in its elastic approximation is much simpler than the full model and does not depend on the physical properties on the reflecting material. Besides that, the energy loss can be treated, in first approximation, as constant, and for this reason we use the elastic Remizovich approximation to model the angular distribution of protons after reflection at glancing angles. In the present simulations we added to the Remizovich scattering a  $3\pm0.7$  keV energy loss, as in [15]. This is the most probable energy loss found in past laboratory measurements of grazing proton scattering on different surfaces in the 30 - 710 keV energy range [19].

The elastic approximation of the Remizovich solution takes the form:

$$W_{el}(\Psi, \chi) = \frac{1}{12\pi^2 \Psi^{1/2}} \left[ \frac{\omega^4}{1 + \omega^2} + \omega^3 \arctan(\omega) \right], \tag{2}$$

$$\omega = \left\{ \frac{3\Psi}{\Psi^2 - \Psi + 1 + (\chi/2)^2} \right\}^{1/2}.$$
 (3)

# 3.2 Firsov model

The elastic probability distribution of Eq. 2 integrated all over the azimuthal angles becomes the formulation obtained by Firsov to describe the reflection of fast ions from a dense medium at glancing angles:

$$W(\Psi) = \frac{3}{2\pi} \frac{\Psi^{3/2}}{1 + \Psi^3}.$$
 (4)

The Firsov formula was used in [15] in the assumption that all protons are scattered at  $\phi = 0$ , i.e. continue their path along the x-axis of Fig. 1 (right panel). In order to compare our results with past simulations we also implemented Eq. 4 with  $\phi = \phi_0 = 0$ . We will generally refer to this implementation as Firsov.

#### 3.3 Single and multiple scattering model

In the single or multiple Coulomb scattering, when a charged particle traverses a medium, it undergoes one or more elastic scatterings due to Coulomb interactions with the electron field of the nuclei, as described by the Rutherford cross section. For grazing incident angles, protons can interact with the the nuclei at the material edge and escape after one or more interactions, with the effect of being scattered by the target, often with an enhanced deflection angle with respect to the incident one if multiple scattering is involved.

Multiple scattering was the first model to be proposed as the one responsible for soft proton funneling by X-ray optics.

## 4 Geant4 simulation set-up

The Geant4 [20–22] Monte Carlo toolkit is a C++ based particle transport code, initially developed by CERN for the simulation of high energy experiments at particle accelerators and then extended to lower energy ranges, i.e. the X and Gamma-ray domain. Geant4 has become the standard tool used by many space agencies (e.g. ESA) in the simulation of the background and instrument performance of all major X-ray space telescopes (e.g., Chandra, XMM-Newton, Suzaku, Athena, eROSITA).

BoGEMMS (Bologna Geant4 Multi-Mission simulator) is a Geant4-based simulation tool developed at the INAF/IASF Bologna [23, 24] for the evaluation of the scientific performance (e.g. background spectra, effective area) of X-ray and Gamma-ray space missions. It allows to interactively set the geometrical and physical parameters recording the interactions in FITS and Root format output files and filtering the output as a real observation in space, to finally produce the background detected count rate and spectra. The BoGEMMS framework and the Geant4 release 10.2 are used throughout the activity presented in this paper, with a selection cut for all volumes of 1 nm. Unless otherwise specified, we use Geant4 reference physics list QGSP\_BERT\_HP with the opt3 electromagnetic physics list on top. Given the involved energies, the simulation does not use the latter list, but it is present to check that the Geant4 simulation correctly selects the new proton physics. Since the validation of proton scattering in Geant4 is achieved by comparing the simulation with the experiment of [16], the mass model of the reflecting surface used for both the verification and validation tests (see Figure 1, left panel) approximates the eRosita X-ray shell sample used in the experiment: a planar slab of Nickel, 270  $\mu m$  thick, coated by 50 nm of Gold.

A new physics list G4SoftProtonPhysics is added for the handling of the two new models describing the scattering of protons at grazing angles: the Remizovich elastic approximation and the Firsov azimuthal elastic integration. Thanks to the BoGEMMS configuration framework, the user can set at run-time the proton energy and incident angular range where to apply the new models. Models can also be combined and used in the same simulation using different energy or incident angle ranges. The probability for the proton to undergo a Firsov or Remizovich interaction is arbitrary set to 100%, if the proton matches the energy and angle range of applicability. This assumption is due to the goal of the present study, i.e. low energy and low angle proton scattering.

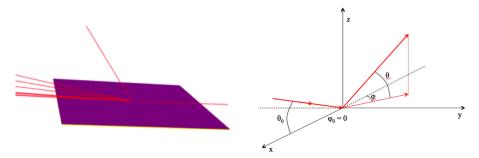


Fig. 1 Left panel: The eRosita Au-coated Ni shell portion is approximated by two planar slabs. The red lines show the proton trajectory before (right side) and after (left side) the interaction. Right panel: The polar  $\theta$  and azimuthal  $\phi$  angle definition used for the Remizovich and Firsov formulation, with the reflecting surface placed in the x-y plane and the proton trajectory highlighted in red.

#### 5 Physics implementation and verification

Among the many physical processes that have been proposed to describe the scattering of soft protons by X-ray optics, only the Coulomb single scattering model is currently available in the official Geant4 toolkit [25]. Geant4 simulations of the soft proton scattering by the XMM-Newton X-ray optics were updated by implementing the Firsov angular scattering distribution on top of the Geant4 version 9.1, with the addition of a constant small energy loss [15]. The two new Geant4 physics classes developed for this purpose, G4FirsovSurface and G4FirsovScattering, were not included in the official release of the Geant4 toolkit.

We define as  $\theta_0$  and  $\phi_0$  the incident polar and azimuthal angles with respect to the reflecting surface and as  $\theta$  and  $\phi$  the polar and azimuthal angles after the proton scattering. For each proton reaching the surface at glancing angle, the x-axis of the Cartesian reference system is placed along the proton trajectory, so that  $\phi_0$  is always zero. Figure 2 shows the angular system used to describe the reflection, with the reflecting material placed in the x-y plane and the proton trajectory highlighted in red. Following the formalism of [17], we introduce the dimensionless variables  $\Psi = \theta/\theta_0$ ,  $\chi = \phi/\theta_0$  to express the proton angular distribution.

Given the simple slab geometry, in both Remizovich and Firsov model implementations no check is inserted on the reflecting surface material, the only requirement for the soft proton scattering to be activated is to match the incident energy and angles. The dependence on the material properties (e.g. density, atomic number) will be inserted in future releases of the physics classes.

## 5.1 Firsov

The range of possible scattering polar angles, from  $0^{\circ}$  to  $90^{\circ}$ , is divided in a limited number of discrete possible values, and the resulting proton angle after scattering is randomly picked up among the list of discrete bins following the analytical probability distribution. The number of bins used to divide the scattering angle range has a direct impact on the resolution of the scattering angles. Since our final goal is to reproduce the experimental results of [16], the uncertainty level

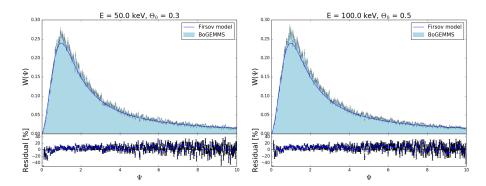


Fig. 2 Comparison between the Firsov angular distribution  $W(\Psi)$  of scattered protons obtained by the BoGEMMS simulation (light blue area) and the analytical model (dark blue line) for an incident energy and polar angle of (50 keV and 0.3°, left panel) and (100 keV and 0.5°, right panel).

in the scattering angle distribution must be lower than the angular dimension of the proton detection area used in the laboratory measurements. In this case, a 1.2 mm side aperture at a focal distance of 933 mm from the center of the X-ray shell translates into an angular resolution of  $\sim 0.07^\circ$ , meaning at least 1000-2000 discrete bins of the polar angle range. For the present simulation, we use a value of 10000.

The simulation set-up for the verification phase consists of a proton point source in correspondence of the selected incident polar angle  $\theta_0$ . Although no dependence is expected on the initial proton energy, we use two different initial energies (50 and 100 keV) to check the correct setting, at run-time, of the proton scattering range of applicability. The resulting probability distribution W as a function of the dimensionless polar variable  $\Psi$  is shown in Fig. 2, for an incident angle of 0.3° (left panel) and 0.5° (right panel). We are able to reproduce the Firsov distribution within  $\sim 10\%$  of the analytical values.

# 5.2 Remizovich

The Remizovich model describes the distribution in both  $\theta$  and  $\phi$ , or  $\Psi$  and  $\chi$ , of the scattered protons. The Geant4 implementation requires dividing in discrete bins both the polar and azimuthal ranges of the scattering angle, from  $0^{\circ}$  to  $90^{\circ}$  and in the  $\pm 90^{\circ}$  range respectively. As described in the previous section for the Firsov case, the number of discrete bins, 3000 for the present simulations, is decided according to the required angular distribution. The binned probability distribution, in the  $\Psi - \chi$  parameter space, obtained from the BoGEMMS simulation and the analytical model is plotted in Fig. 3, with the color bar showing the value of the binned distribution  $W(\Psi,\chi)$ . Fig. 4 shows the linear comparison between the BoGEMMS and the Remizovich model obtained by integrating over a range of scattering azimuthal angles  $\psi = \pm 20^{\circ}$  (left panel) and polar angles  $0^{\circ} < \theta < 40^{\circ}$  (right panel). The model is reproduced within a maximum uncertainty of  $\sim 20\%$ . The error bars are  $1\sigma$  Poisson fluctuations of the number of detected events for each bin and depend on the total number of emitted protons.

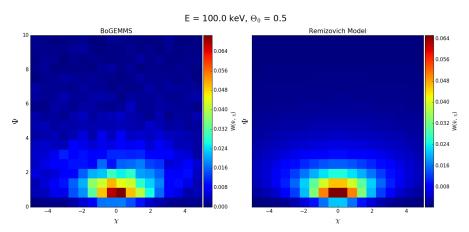


Fig. 3 Comparison between the bi-dimensional Remizovich angular distribution  $W(\Psi,\chi)$  of scattered protons obtained by the BoGEMMS simulation (left panel) and the analytical model (right panel) for an incident energy and polar angle of 100 keV and 0.5°.

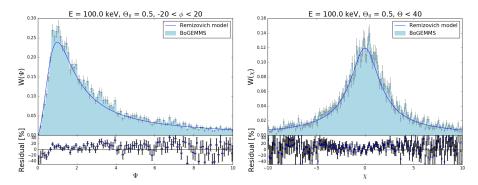


Fig. 4 Linear comparison of the integrated Remizovich distribution, for a polar incident angle of 0.5°, by integrating the azimuthal angle range  $\psi=\pm20^\circ$  (left panel) and the polar angle  $0^\circ<\theta<40^\circ$  (right panel).

The Remizovich implementation in the Geant4 code has been compared to the ray-tracing simulation of [26] for an independent verification. Contrary to the Geant4 particle transport code, a ray-tracing Monte Carlo simulator follows each particle from the mirror to the focal plane by applying the reflection model of the optics design. The present ray-tracing code, used in the design of the MECS on board of BebboSAX [27] and for the calibration of the Swift XRT effective area [28], has been modified to account for the reflection of soft protons by the ATHENA pore optics. As shown in Fig. 5, the scattering efficiency (see Sec. 6.1 for a detailed description) predicted by the two simulators using the Remizovich solution is in very good agreement, confirming the proper construction of the process in Geant4.

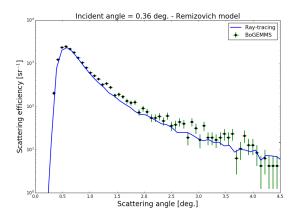


Fig. 5 Comparison of Remizovich induced scattering efficiency obtained with the two independent Geant4 (crosses) and ray-tracing (continuous line) simulations.

#### 5.3 Single and multiple scattering

In Geant4 10.2 there are two models describing the scattering of protons [29], applied here to grazing angle interactions: the Urban model of multiple scattering (G4UrbanMscModel) used in the electromagnetic physics list opt3 and a combination of multiple scattering G4WentzelVIMscModel and single scattering model G4eCoulombScatteringModel used by default and for the opt4 physics list. These combinations of models are working simultaneously: one of the two models is called depending on the step and the relative probability. Small scattering angles are sampled by the multiple scattering model, large scattering angles by single scattering. These model are coherent, because they are using the same elastic cross section of Wentzel [30]. In the G4EmStandardPhysicsSS physics list instead only the single scattering model is defined, for any scattering angle.

The Urban model uses empirical parameterisations to sample large scattering angles, which may not be accurate for all cases. It requires validation for each particular setup and often extra step limitations to get better agreement with the data. The WentzelVI multiple scattering model is more accurate but also requires checks of the optimal step limit for concrete uses cases. The single scattering model may be used out of the box and does not need a special tuning of step limits.

These three types of physics lists simulate different trajectories of grazing protons in the absorber, and for this reason energy losses computed by Geant4 ionisation models can be different.

# 6 Comparison with real data

The uncertainty in the incident proton angle is simulated using as proton source a beam profile with a standard deviation equal to the angle error, as shown in Fig. 6. In addition to the Firsov and Remizovich formula, the proton scattering is simulated using the following available physics processes:

- Multiple Coulomb scattering, provided by the default *G4EmStandardPhysics\_option3* electromagnetic physics list (opt3);

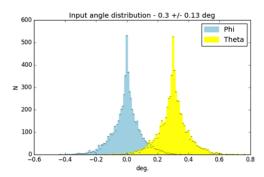


Fig. 6 The polar, in yellow, and azimuthal, in light blue, angular distribution of the proton beam profile at an incident angle of  $0.3^{\circ}$  and a standard deviation of  $0.13^{\circ}$ .

- Multiple Coulomb scattering, provided by the G4EmStandardPhysics\_option4 electromagnetic physics list (opt4);
- Single Coulomb scattering, provided by the G4EmStandardPhysicsSS electromagnetic physics list.

All processes have been used with the standard settings provided by the Geant4 10.2 release.

#### 6.1 Scattering efficiency

Since the scattering angular distribution is not isotropic, it is important in the simulation to collect the protons in the same solid angle used in the experiment. The results presented here are obtained by selecting only the scattered protons within the azimuthal range  $\phi = \pm 0.037^{\circ}$ , the detection aperture subtended angle, and dividing the polar range in discrete bins in order to obtain, for each bin, the aperture solid angle  $\Omega$ . The number of protons in each bin is  $N_{\rm det}$  of Eq. 1. Being both Firsov and Remizovich models elastic - the resulting scattering efficiency does not depend on the initial proton energy - we compare the output of the simulation with respect to experimental results for the three proton energies. The scattering efficiency for an incident proton energy of 250 keV and incident angle of 0.36° is shown in Fig. 7. The Firsov formula implemented with  $\phi = 0$  has the effect of focusing all protons in the detection area with the results of overestimating by more than an order of magnitude the scattering efficiency. Multiple scattering, for both opt3 and opt4, causes at 250 keV scattering angles larger than the values obtained in the experiment: as shown in Fig. 9 (left panel), the scattering polar angle peaks in the  $10^{\circ} - 20^{\circ}$  range.

If small incident angles are considered, the Remizovich and the single scattering (SS) both well reproduce the proton angular distribution at scattering angles higher than 1°. Near the specular reflection ( $\sim 0.7^{\circ}$ ) both models give higher values, with the SS inducing a scattering efficiency closer to the experimental data (see App. A for a comparison at all tested incident angles).

At 500 keV and  $\theta_0 = 0.33^{\circ}$  (Fig. 8, left panel), MSC protons are also visible below a scattering angle of 4.5°, but with lower efficiencies at very small angles. The *opt3* 

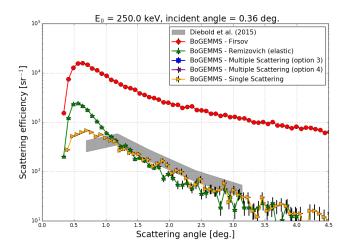


Fig. 7 Scattering efficiency at  $E_0 = 250 \text{ keV}$  for an incident angle of  $0.36^{\circ}$ . The grey area shows the experimental results. Multiple scattering, both in *opt3* and *opt4*, results in higher scattering angles.

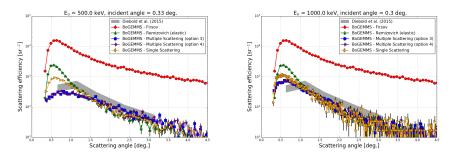
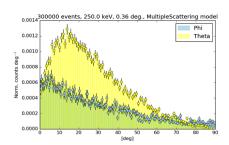


Fig. 8 Left panel: Scattering efficiency at  $E_0 = 500$  keV for an incident angle of  $0.33^{\circ}$ . Right panel: Scattering efficiency at  $E_0 = 1000$  keV for an incident angle of  $0.3^{\circ}$ .

and opt4 lists result in the same distribution. SS efficiency rises at small angles for  $E_0=500~\rm keV$  and is consistent among the entire range of experimental scattering angles. If protons of 1000 keV are emitted (Fig. 8, right panel), SS and MSC give similar results, with about a factor 2 of difference at the specular scattering angle. All models, except for the Firsov formula, generate consistent angular distributions for scattering angles higher than 1°. Considering the three incident energies, 250, 500 and 1000 keV, the SS is the model that results in more similar results to the real data set. In Fig. 9 (right panel) we compare the SS induced proton angular distribution for two extreme energy values, 1000 and 50 keV. Above  $\sim 1.2^\circ$ , for the present detection geometry, the proton energy has no effect on the scattering efficiency, while it becomes important at very small scattering angles, with about a factor 4 of difference between the two cases.



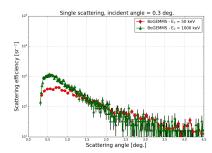


Fig. 9 Left panel: The angular distribution in  $\theta$  (yellow points) and  $\phi$  (light blue points) of scattered protons if multiple scattering is used as physics interaction. Right panel: Single scattering efficiency at  $E_0 = 50$  and 1000 keV for an incident angle of  $0.3^{\circ}$ .

### 6.2 Energy losses

The measured energy loss is obtained by Gaussian fits of the energy distribution of both the incident and scattered protons. In our case the statistics for each bin of the scattering angle is not enough to produce a fit and the energy loss is given by the mean of the proton energies, for each bin, subtracted by the incident energy (error bars are the standard deviation of the energy distribution). The experimental results, that range from  $\langle E_{Loss} \rangle = 13 \text{ keV } (\theta_0 = 0.69^\circ)$  at  $E_0 = 250 \text{ keV }$  to 54 keV ( $\theta_0 = 3.09^\circ$ ) at  $E_0 = 1000 \text{ keV}$ , give higher energy losses than the few keVs expected from [19]. From these findings, the percentage of energy lost in each scattering seems to be constant, with a value of  $\sim 5\%$  with respect to the initial energy. Remizovich and coauthors [17] find a constant behaviour in the percentage of energy lost in the scattering similar to what obtained in the present measurements. Since the Remizovich model is implemented in the elastic approximation, we only compare the simulation using the inelastic SS and MSC scattering interactions.

At 250 keV and for  $\theta_0 = 0.36^{\circ}$  (see Fig. 10), SS gives energy losses less than 1 keV at the specular reflection angles, and up to  $\sim 10$  keV for larger scattering angles, about 10 times less than the experimental data. The same behavior is obtained at higher incident angles (see Fig. 15 of App. B). At 500 and 1000 keV (see Fig. 11), SS and MSC give similar results at large scattering angle, as seen for the scattering efficiency, but in the specular range the multiple scattering induces higher, > 10 keV, energy losses, close to the real data points. No difference is observed between the electromagnetic *opt3* and *opt4*. For larger incident angles (see Figure 72 and Figure 73 in Appendix) the multiple scattering results in energy losses well consistent with the observed ones if the proton energy is in the 500-1000 keV band.

## 6.3 Code optimization issues

In the present Geant4 implementation, the Remizovich probability distribution, which depends on the incident proton angle, is computed at each interaction. This approach is feasible for the Firsov model, where only the polar angle is randomly

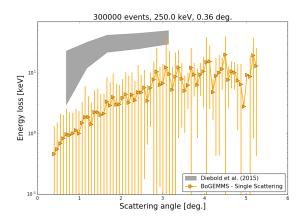


Fig. 10 The proton energy loss as a function of the scattering angle at  $E_0 = 250 \text{ keV}$  for an incident angle of  $0.36^{\circ}$ .

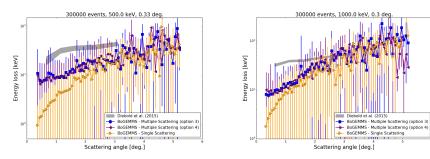


Fig. 11 Left panel: The proton energy loss as a function of the scattering angle at  $E_0 = 500$  keV for an incident angle of  $0.33^{\circ}$  Right panel: The proton energy loss as a function of the scattering angle at  $E_0 = 1000$  keV for an incident angle of  $0.3^{\circ}$  deg.

 $\textbf{Table 1} \ \, \textbf{CPU} \ \, \textbf{time, normalized to the multiple scattering (opt 3) time, required to run the simulation for the tested physics models.$ 

	MSC(opt3)	MSC (opt4)	SS	Remizovich	Firsov
CPU time	1	1.1	21.3	366.7	3.1

generated, but not for the Remizovich model, that is a bivariate distribution in  $\phi$  and  $\theta$ , with a CPU processing time higher than a factor 100 with respect to the integrated Firsov version at the highest energies. Table 1 compares the simulation CPU time for all the tested models in units of time required by multiple scattering (emstandard\_option3). The performance test refers to a run of  $10^4$  protons at 250 keV for an incident angle of 0.36 degrees.

Thanks to the very simple mass model used in the proton scattering test, we were able to perform the present simulations by a dedicated fine-tuning of the angle resolution (see Sec. 5.1) and the optimization of the Geant4 algorithm. The latter produced a gain of  $\sim 50\%$  in CPU time. In case of future releases of the code for its use in simulation campaigns of the background of X-ray space telescopes, as

the case of ATHENA, where complex geometries and high statistics are required, the optimization of the Remizovich Geant4 implementation is mandatory. Besides the use of the Geant4 multi-threading option already built in the standard Geant4 10.2 release, possible solutions could be the use of for loop parallelization APIs (e.g. openMP<sup>1</sup>) or loading the probability distribution from an external physics database.

#### 7 Summary

The most recent laboratory measurements of [16], testing the angular and energy distribution of protons scattering by X-ray mirror shells at glancing angles, are the only available data set for the physics assessment of the interaction behind the soft proton scattering in X-ray space telescopes. After the implementation and verification of two new physics models (the elastic Remizovich model and the azimuthally focused Firsov model), we simulate the experimental set-up to find out which model, among the new solutions and the ones already built in the standard Geant4 library (single and multiple scattering), better reproduce the observation. Our results can be summarized as follows:

- if we consider the proton distribution at 250 keV, Remizovich and SS are well consistent with the experimental scattering efficiency except for very small,
   1°, scattering angles, where higher efficiencies are found in the simulation;
- the SS induced scattering efficiency at small angles is the closest to the observation, but the energy losses are a factor 10 less than the experimental ones;
- MSC can not reproduce the angular distribution of protons at 250 keV, and it
  is not feasible to simulate soft proton funnelling by X-ray optics;
- at large scattering angles, the SS and Remizovich solutions give consistent scattering efficiencies;
- because of the large spread in the azimuthal scattering angle, the use of the Firsov model with  $\phi = 0$  overestimates more than 10 times the real distribution;
- no differences are found between the electromagnetic opt3 and opt4 list MSC settings.

## 8 Conclusions and remarks

Since the Remizovich model in its approximated form has proven a general consistency with the measurements, the present Remizovich Geant4 implementation will not only be included in the  $Space\ Physics\ list$  of the ATHENA Radiation Environment Models and X-Ray Background Effects Simulators (AREMBES), but it is currently being used as basis for the development of the official Geant4 Remizovich classes to be included in the next release of the Geant4 toolkit. This activity is carried on under the responsibility of Geant4 collaboration members within the  $Low\ Energy\ working\ group\ ^2$ . The implementation in Geant4 of the Remizovich formula for inelastic scattering, predicting energy losses close, within a factor 2-4, to what obtained in the latest measurements, should provide a better description

<sup>&</sup>lt;sup>1</sup> http://openmp.org/wp/

 $<sup>^2\</sup> https://twiki.cern.ch/twiki/bin/view/Geant4/LowEnergyElectromagneticPhysicsWorkingGroup (Control of the Control of Control of$ 

of the experimental results and it is planned in the future activities.

Despite the obtained results represent a first step toward the development of a Geant4 grazing angle low energy proton scattering model dedicated to X-ray optics on board space missions, it must be pointed out that the lowest proton energy obtained in this experimental facility, 250 keV, is above the range of interest of the protons that induce background in X-ray missions like Chandra and XMM-Newton, where protons below 100 keV are the ones that mostly deposit energies inside the detectors sensitivity band, inducing a poorly reproducible background component.

Because of (i) the lack of data at very low scattering angles, (ii) the lower energy losses obtained with respect to the experiment, (iii) and the high energy of the incoming protons, it is not yet possible to advertise any of the tested models as the most accurate one for the simulation of the scattering of low energy protons experienced by ATHENA X-ray optics. However, further developments are both ongoing on the software side and studies planned as new data will be available in the near future, focused at lower proton energies and lower scattering angles in order to definitively select one of those models to represent the proton scattering behavior in this regime.

Acknowledgements This work is supported by the ESA Contract No 4000116655/16/NL/BW. The AHEAD project (grant agreement n. 654215) which is part of the EU-H2020 programm is acknowledged for partial support.

#### References

- 1. M.C. Weisskopf, et al., (2000), Proc. SPIE, vol. 4012, pp. 2–16
- 2. F. Jansen, et al., A&A 365, L1 (2001). DOI 10.1051/0004-6361:20000036
- 3. R. Willingale, XRT-LUX-RE-011/1 Technical Report (2000)
- 4. G.P. Garmire, et al., (2003), Proc. SPIE, vol. 4851, pp. 28-44. DOI 10.1117/12.461599
- G.E. Villa, et al., (1996), Proc. SPIE, vol. 2808, pp. 402–413
   J.A. Carter, A.M. Read, A&A 464, 1155 (2007). DOI 10.1051/0004-6361:20065882
- 7. A. De Luca, S. Molendi, A&A 419, 837 (2004). DOI 10.1051/0004-6361:20034421
- 8. D.H. Lumb, et al., A&A 389, 93 (2002). DOI 10.1051/0004-6361:20020531
- 9. D.N. Burrows, et al., Space Science Review 120, 165 (2005). DOI 10.1007/ s11214-005-5097-2
- 10. K. Mitsuda, et al., PASJ 59, 1 (2007). DOI 10.1093/pasj/59.sp1.S1
- 11. U.G. Briel, et al., (2000), Proc. SPIE, vol. 4012, pp. 154-164
- 12. K. Nandra, et al., ArXiv e-prints (2013)
- 13. P. Predehl, et al., (2014), Proc. SPIE, vol. 9144, p. 91441T. DOI 10.1117/12.2055426
- 14. D. Spiga, et al., (2008), Proc. SPIE, vol. 7011, p. 70112Y. DOI 10.1117/12.789917
- 15. F. Lei, et al., IEEE Trans. Nucl. Sci.  ${\bf 51}(6),\,3408$  (2004)
- DOI 10.1007/ 16. S. Diebold, et al., Experimental Astronomy 39, 343 (2015). s10686-015-9451-4
- 17. V.S. Remizovich, M.I. Ryazanov, I.S. Tilinin, Sov. J. Exp. Th. Phys. 52, 225 (1980)
- 18. Mashkova, E. S., et al., Rad. Effects p. 85 (1983)
- 19. K. Kimura, M. Hasegawa, M.H. Mannami, Phys. Rev. B 36, 7 (1987). DOI 10.1103/ PhysRevB.36.7
- 20. S. Agostinelli, et al., NIM A 506, 250 (2003)
- 21. J. Allison, et al., IEEE Trans. Nucl. Sci.  ${\bf 53}(1),\,270$  (2006)
- 22. J. Allison, et al., NIM A 835, 186 (2016). DOI 10.1016/j.nima.2016.06.125
- 23. A. Bulgarelli, et al., (2012), Proc. SPIE, vol. 8453, p. 845335. DOI 10.1117/12.926065
- 24. V. Fioretti, et al., (2014), Proc. SPIE, vol. 9144, p. 91443N. DOI 10.1117/12.2056442
- 25. R. Nartallo, et al., IEEE Trans. Nucl. Sci. 48(6), 1815 (2001)
- 26. T. Mineo, et al., INAF Tech. Report INAF-XIFU-TM-2015-1 (2015)

- G. Conti, et al., (1994), Proc. SPIE, vol. 2279, pp. 101–109
   G. Cusumano, et al., (2006), American Institute of Physics Conference Series, vol. 836, pp. 664–667. DOI 10.1063/1.2207972
   V.N. Ivanchenko, et al., (2010), J. Phys. Conf. Ser., vol. 219, p. 032045. DOI 10.1088/1742-6596/219/3/032045
   G. Wentzel G, Z. Phys. 40, 590 (1927)

# A Scattering efficiency

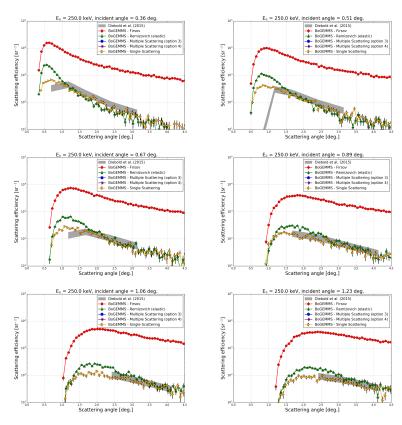


Fig. 12 Scattering efficiency at  $E_0=250~\rm keV$  for an incident angle ranging from  $0.36^\circ$  to  $1.23^\circ.$ 

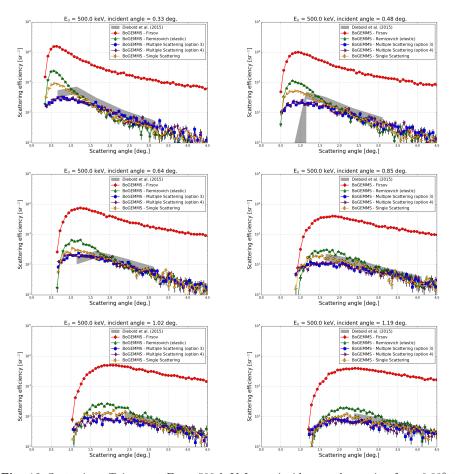


Fig. 13 Scattering efficiency at  $E_0=500~{\rm keV}$  for an incident angle ranging from  $0.33^\circ$  to  $1.19^\circ.$ 

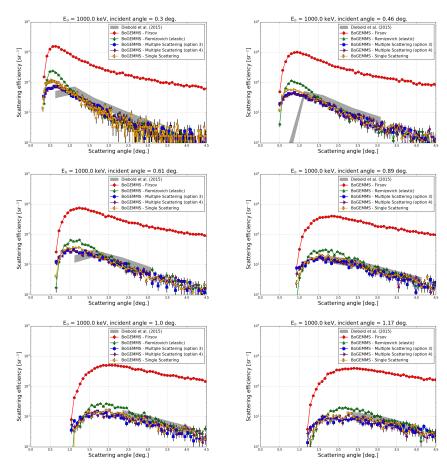


Fig. 14 Scattering efficiency at  $E_0=1000~\rm keV$  for an incident angle ranging from  $0.3^\circ$  to  $1.17^\circ$ .

# B Energy losses

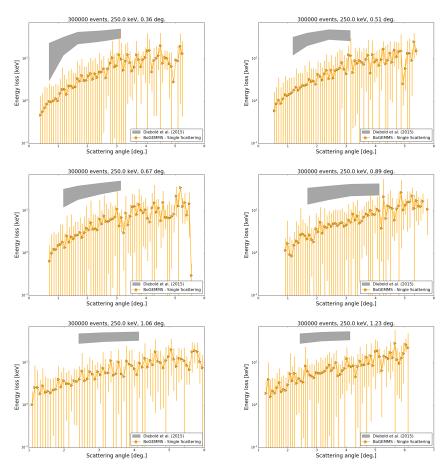


Fig. 15 The proton energy loss as a function of the scattering angle at  $E_0=250~\rm keV$  in the  $0.36^\circ-1.23^\circ$  incident angle range.

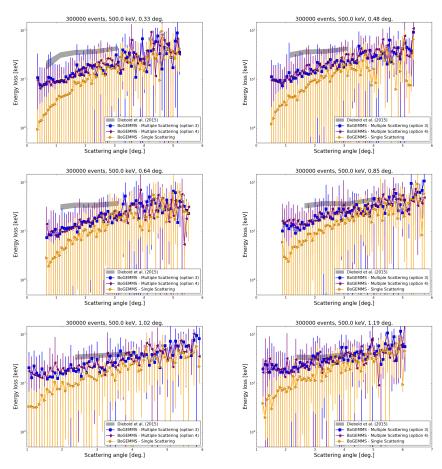


Fig. 16 The proton energy loss as a function of the scattering angle at  $E_0=500$  keV in the  $0.33^\circ-1.19^\circ$  incident angle range.

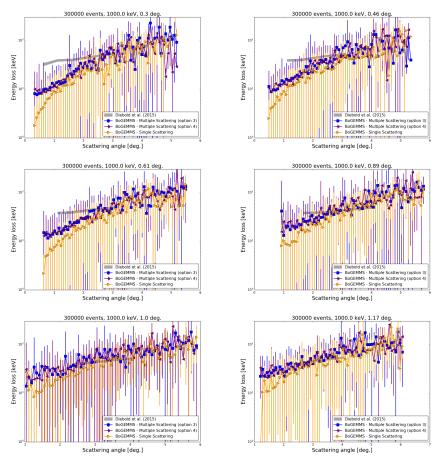


Fig. 17 The proton energy loss as a function of the scattering angle at  $E_0=1000~\rm keV$  in the  $0.3^\circ-1.17^\circ$  incident angle range.