

Figure 12. Time-series analysis of PG 1351+640. The black dotted and solid lines in panels (d)–(h) are the original and cleaned (e.g., Fe II, He II, narrow H β , and [O III] λ 4959,5007) mean spectra. The meanings of the other panels, lines, and histograms are the same as in Figure 3.

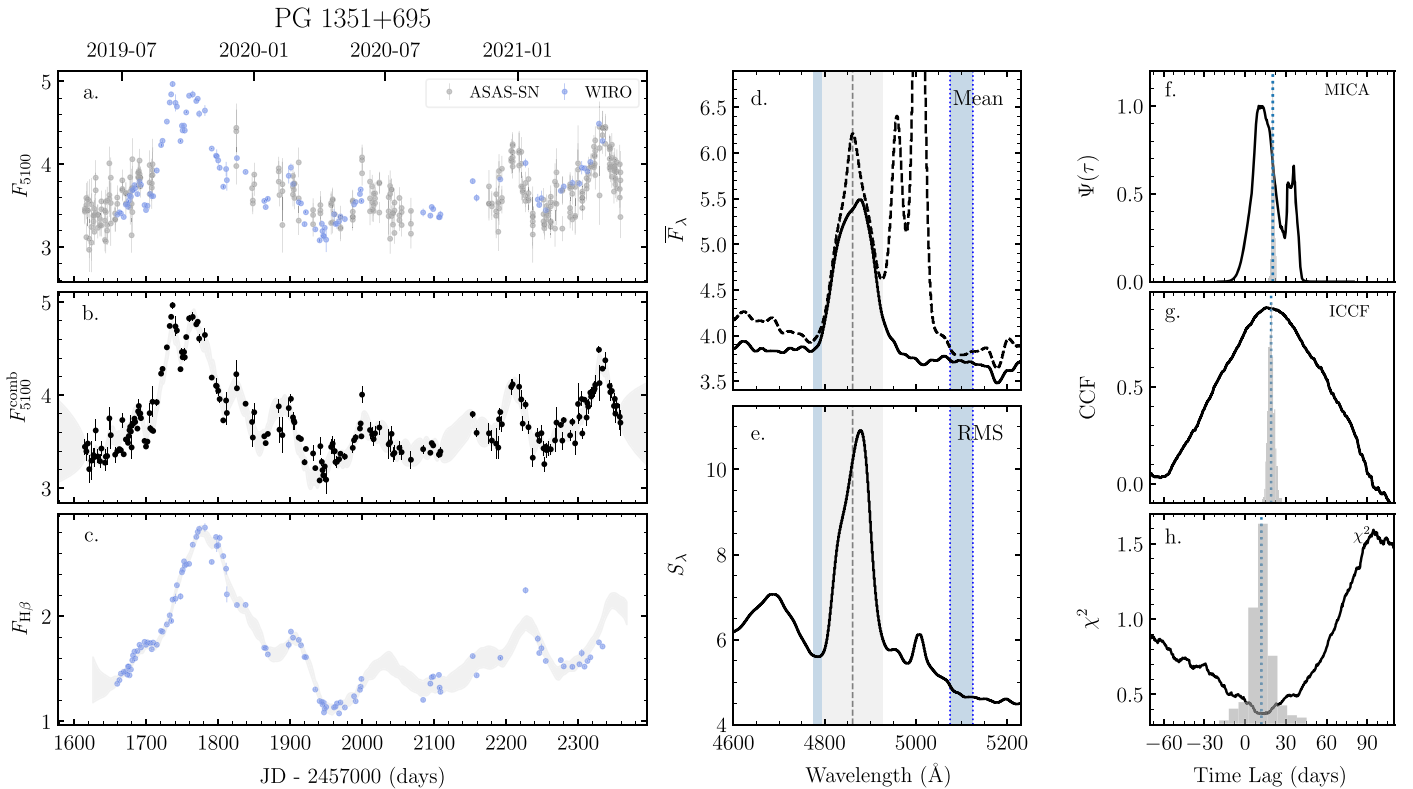


Figure 13. Time-series analysis of PG 1351+695. The meanings of the panels, lines, and histograms are the same as in Figure 2.

The FWHM and $\sigma_{H\beta}$ from the mean and rms spectra are provided in Table 6. The H β signal in the rms spectra of some objects is too weak for the line width measurement [PG 0049 +171 (Season 4), PG 0947+396 (Seasons 2 and 4), PG 1202

+281 (Seasons 1–4), PG 1351+640, PG 1501+106 (Season 2)]. We measured the FWHM and $\sigma_{H\beta}$ of H β from the aforementioned continuum-cleaned rms spectra for all seasons (see Section 3.1).

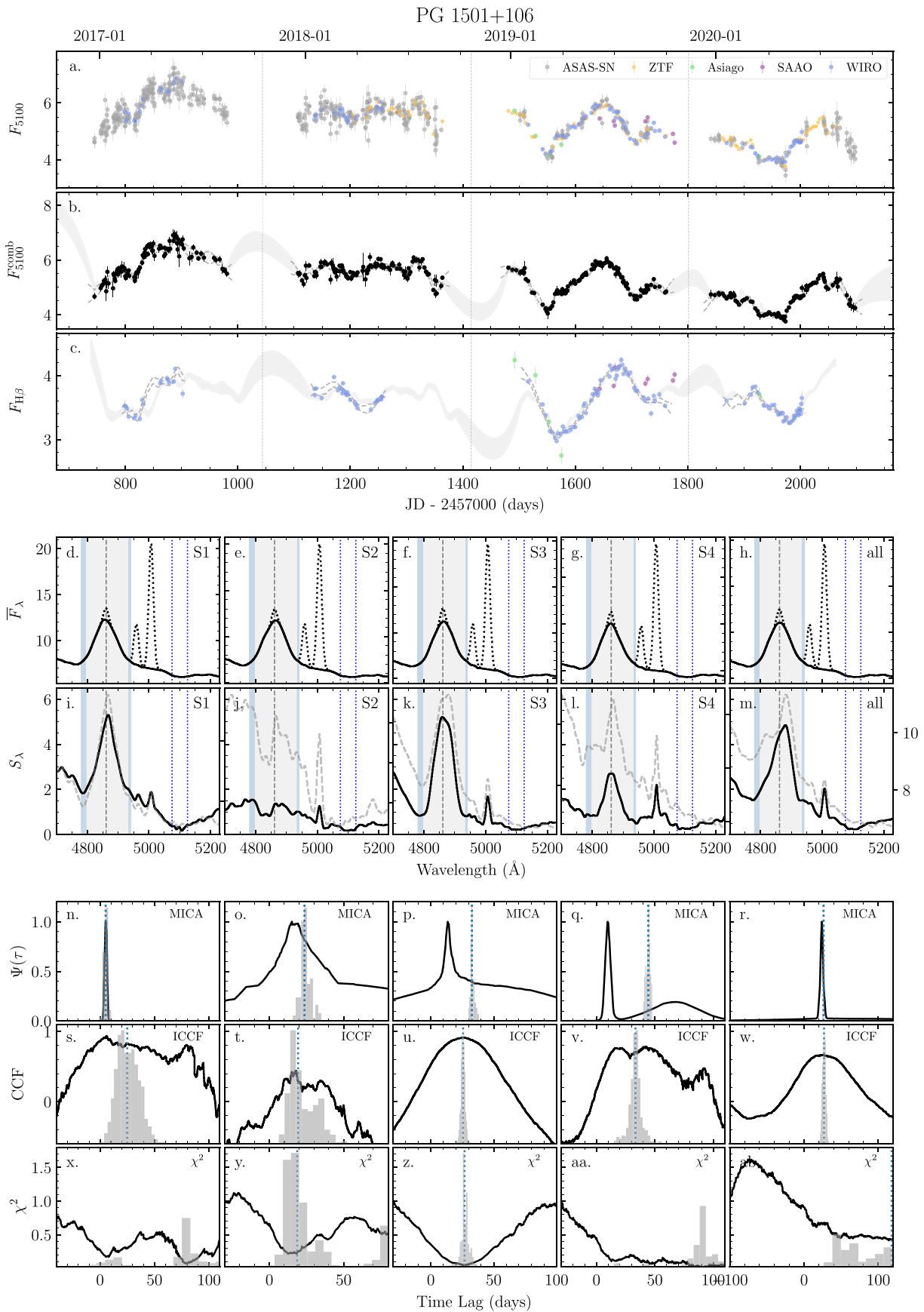


Figure 14. Time-series analysis of PG 1501+106. The meanings of the panels, lines, and histograms are the same as in Figure 3.