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The electronics of the High-Energy Particle Detector on board the CSES satellite

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Abstract

CSES (China Seismo-Electromagnetic Satellite) is a Chinese-Italian scientific space mission dedicated to monitor the variations of the main parameters of the topside ionosphere (electric and magnetic fields, plasma parameters, charge particles fluxes) caused by either natural emitters - especially earthquakes - or artificial ones.

The CSES satellite was successfully launched from the Jiuquan Satellite Launch Center located in the west of Inner Mongolia on February 2nd, 2018, and it is now orbiting under nominal conditions. The expected mission lifetime amounts to 5 years. CSES is the first element of a multi-satellite monitoring system; several satellites are scheduled for the next few years.

The High Energy Particle Detector (HEPD) is the main contribution of the Italian collaboration to the mission. It was designed and built in order to detect electrons in the energy range between 3 and 100 MeV, protons between 30 and 200 MeV, and light nuclei in the MeV energy window.

The electronics of the detector was designed following stringent requirements on mechanical and thermal stability, power consumption, radiation hardness and double redundancy. The system successfully went through the space qualification tests. In this paper, we describe the HEPD electronics, the space qualification tests performed before launch, and the in-flight performance of the detector.

Keywords: Detector techniques for Cosmology and Astroparticle Physics, Satellite experiment, electronics

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1. Introduction

CSES (China Seismo-Electromagnetic Satellite) [1] is a Chinese-Italian scientific space mission devoted to the monitoring of electromagnetic fields and waves, plasma and particles perturbations of the atmosphere, ionosphere and magnetosphere induced by natural sources and anthropogenic emitters, and to investigate their correlations with the occurrence of seismic events. In general, CSES is intended to study the structure and dynamics of the topside ionosphere, the coupling mechanisms with the lower

and higher plasma layers, and the temporal variations of the geomagnetic field, under quiet and disturbed conditions. Solar-terrestrial interactions and solar physics phenomena, namely Coronal Mass Ejections (CMEs), solar flares and cosmic ray solar modulation, will also be studied.

Eight payloads are installed on board the CSES satellite for the measurement of electromagnetic field components, plasma parameters and energetic particles, as well the as the X-ray flux, namely: a High-Precision Magnetometer (HPM), a Search-Coil Magnetometer (SCM), an Electric Field Detector (EFD), a Plasma Analyzer Package (PAP),

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23 a Langmuir Probe (LP), a Global Navigation Satellite Sys-
 24 tem (GNSS) Occultation Receiver, and a set of particle
 25 detectors, the High-Energy Particle Package (HEPP) and
 26 the High-Energy Particle Detector (HEPD). The Italian
 27 contribution to the mission includes the design and con-
 28 struction of HEPD, which is devised to detect electrons in
 29 the energy range between 3 and 100 MeV, protons between
 30 30 and 200 MeV, and light nuclei.

31 The article is organized as follows: after an overview
 32 of the CSES satellite (section 2) and HEPD instrument
 33 (Section 3), Section 4 describes the general architecture
 34 of HEPD electronics and each of its sub-systems in de-
 35 tail. Section 5 summarizes the tests and qualification cam-
 36 paigns carried out on the four models of HEPD before
 37 launch. Section 6 is devoted to the presentation of the in-
 38 orbit performance of HEPD electronics after two years of
 39 flight. Finally, Section 7 presents a summary and conclu-
 40 sions.

41 2. CSES satellite

42 The CSES program plans to put multiple satellites into
 43 orbit to allow concurrent observations of the physical phe-
 44 nomena of interest. Zhangheng-1, the first CSES satellite
 45 (aka CSES-01), has been in orbit since February 2nd, 2018.
 46 The second satellite (Zhangheng-2 or CSES-02) is under
 47 development, and its launch is currently foreseen for early
 48 2022.

49 CSES-01 is a 3-axis stabilized satellite based on the
 50 CAST2000 platform with a mass of 730 kg and a size of
 51 145 cm (Y) × 144 cm (Z) × 143 cm (X) (see figure 1).
 52 It mounts a single deployable solar array and has a peak
 53 of power consumption of about 900 W. Scientific data are
 54 downloaded 6 to 8 times a day for about 500 seconds. The
 55 maximum transmission speed is 120 Mbit/s. CSES-01 is
 56 maintained in a Sun-synchronous orbit at an average alti-
 57 tude of 500 km, an inclination of 97.32°, a period of 94.6
 58 minutes, a revisit time of 5 days, and the LTDN (Local
 59 Time of Descending Node) is at 14:00 PM. The planned
 60 life of the mission is 5 years.

61 CSES-01 comprises the following platform subsystems:
 62 Attitude and Orbit Control (AOC), On-Board Data
 63 Handling (OBDH), Tracking, Telemetry and Command
 64 (TTC), Power Supply (composed of an 80-Ah Li-ion bat-
 65 tery and solar panels), and Thermal Control.

66 The AOC makes use of Earth-oriented 3-axis stabiliza-
 67 tion; attitude sensors (three star trackers, two groups of
 68 gyros, and one digital sun sensor) are used to measure the
 69 attitude, reaction wheel and magnetic torque in order to
 70 maintain the zero-momentum control.

71 To reduce any interference over the scientific payloads
 72 from solar panel rotation or AOC adjustments, two oper-
 73 ating regions are selected:

- 74 • the payload working zone at latitudes between -65°
 75 and $+65^\circ$,

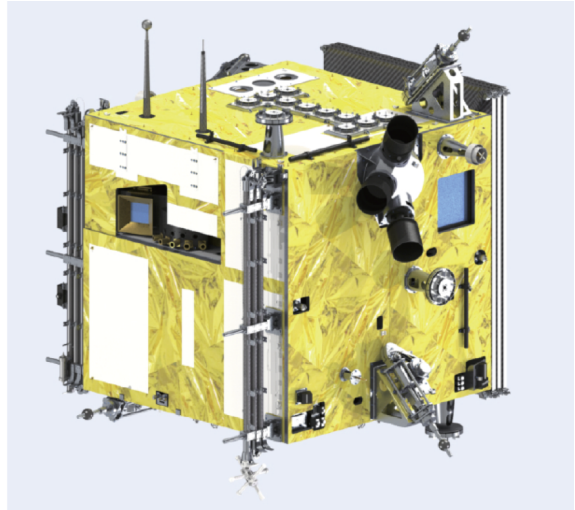


Figure 1: Layout of the CSES satellite with stowed solar panel.

- 76 • the platform adjustment zone at latitudes $> +65^\circ$ or
 77 $< -65^\circ$, where the payloads stop working.

78 When in working zone, each instrument collects data in
 79 two operating modes: “burst mode” and “survey mode”.
 80 The burst mode is usually activated when the satellite
 81 passes over China and regions of the world marked by the
 82 strongest seismic activity. The survey mode is thought for
 83 the remaining areas of the Earth. When in orbit, the X
 84 axis of the satellite is oriented along the velocity vector,
 85 while the Z axis points to nadir. The solar panel located
 86 on one side of the satellite can be rotated around the Y
 87 axis at latitudes $> +65^\circ$ and $< -65^\circ$ in order to optimize
 88 power consumption.

3. The High Energy Particle Detector

The High Energy Particle Detector aims to investigate
 the precipitation of trapped particles induced by atmo-
 spheric electromagnetic (EM) emissions and seismo-EM
 disturbances [2, 3, 4]. HEPD provides good energy and
 angular resolution for electrons (3-100 MeV), protons (30-
 200 MeV), and light nuclei.

The instrument consists of two main active sensors, a
 silicon tracker to measure the impact point and arrival di-
 rection of any impinging particle, and a range calorimeter
 to measure particle energy. In more detail, as shown in
 Figure 2, the instrument consists of:

- two planes of double-side silicon microstrip sensors
 placed on top of the instrument (silicon tracker);
- one layer of segmented plastic scintillator (trigger
 plane);
- 16 plastic scintillator layers, plus a layer of LYSO
 (Lutetium-Yttrium Oxyorthosilicate) inorganic scin-
 tillator crystals (range calorimeter);

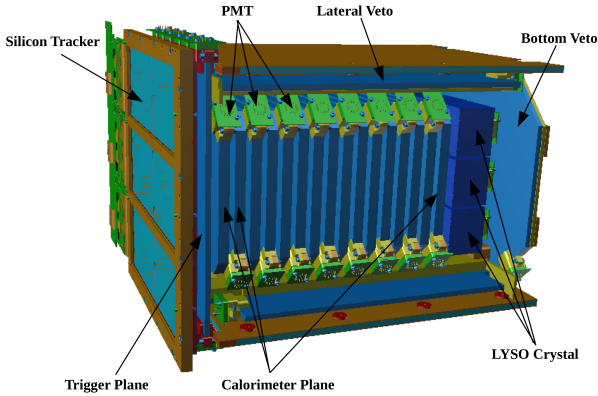


Figure 2: View of the HEPD apparatus: lateral and top panels have been removed for visualization purposes.

- one layer of plastic scintillator (veto bottom).

Additional plastic scintillators cover the four sides of the apparatus, completing the veto system.

Every plastic scintillator is read out by two photomultipliers (PMTs), while the LYSO crystals are read out by one single PMT located to its bottom side. In all cases, the readout is performed by R9880U-210 Hamamatsu PMTs. The HEPD detector is contained within an aluminum box with dimensions $40.36 \times 53.00 \times 38.15 \text{ cm}^3$. The box is housed in the satellite cabin space, which provides the contact surface for heat dissipation. The total instrument mass is about 45 kg, the power budget is 43 W. A detailed description of the HEPD sub-detectors can be found in [5, 6].

4. The electronics of HEPD

4.1. General architecture

HEPD electronics has been designed according to the general requirements imposed on the payload by satellite operations: in this specific case, a power budget of 43 W, a temperature operating range between $-10 \text{ }^\circ\text{C}$ and $+35 \text{ }^\circ\text{C}$, a limit on the maximum amount of data transferable to the ground per day (data budget: 50 Gb/day) and a lifetime longer than 5 years.

The HEPD electronic system is composed of seven subsystems:

- Main Control System;
- Front-end electronics and Analog-to-Digital conversion (ADC) electronics of the Silicon tracker;
- Front-end electronics and Analog-to-Digital conversion electronics of the PMTs (Trigger, Calo and Veto systems);
- Trigger system;
- Data acquisition (DAQ) system;

- Housekeeping system;
- Power system.

The **Main Control system** acts as the main command interface of HEPD with the satellite platform. The bi-directional CAN-bus interface allows to receive both satellite broadcast commands and ground-originated telecommands. It also allows to send command replies and instrument telemetry back. Power-on and power-off sequences, as well as the transition across different HEPD operational modes (SAFE, NOMINAL and STAND-BY), are performed via telecommands received onto the CAN-bus interface.

The Silicon Tracker **Front-end** allows to read out the signals on the silicon microstrips via a hybrid analog and digital electronics. The signals are then sent to the ADCs, and to a Digital Signal Processor that performs the zero suppression and sends digital data to the DAQ system. The PMT front-end and DAQ system similarly digitize signal amplitudes and forward them as digital data to the DAQ system.

The **Trigger System** continuously collects above-threshold digital information from the PMTs, starting digitization of interesting events. It also produces digital data about the triggered signal pattern that are sent to the DAQ system to be included in the scientific datastream.

The **Data Acquisition System**, as previously stated, collects data from the Silicon Tracker, the PMTs and the Trigger system, thus producing the scientific data packets that consist - for each acquired event - of the information on the impact point of the particle in the Silicon tracker and the energy released into the trigger planes, the calorimeter and possibly the veto system. These data are sent to ground using a data link, based on the RS422 protocol, which allows the transmission of the compressed scientific data to the Satellite On Board Data Handler (OBDH).

The **Housekeeping system** monitors the status of the instrument by periodically acquiring the parameters and status register of the subsystems via the internal control link bus (SpaceWire link). The housekeeping data are sent to the satellite and/or the DAQ in order to be integrated into the scientific datastream.

The **Power system** receives the 29 V primary bus from the satellite and provides supply voltages to the other electronic systems, and the bias voltages to the silicon sensors and the PMTs.

The block diagram of HEPD electronics is shown in figure 3.

The HEPD-01 electronics systems are physically hosted in seven boards, all located in a dedicated box connected to the HEPD baseplate (figure 4) apart from the Silicon tracker front-ends, which are located close to the detector. The seven boards are:

- CPU board - it hosts the Main Control system and the Housekeeping system;

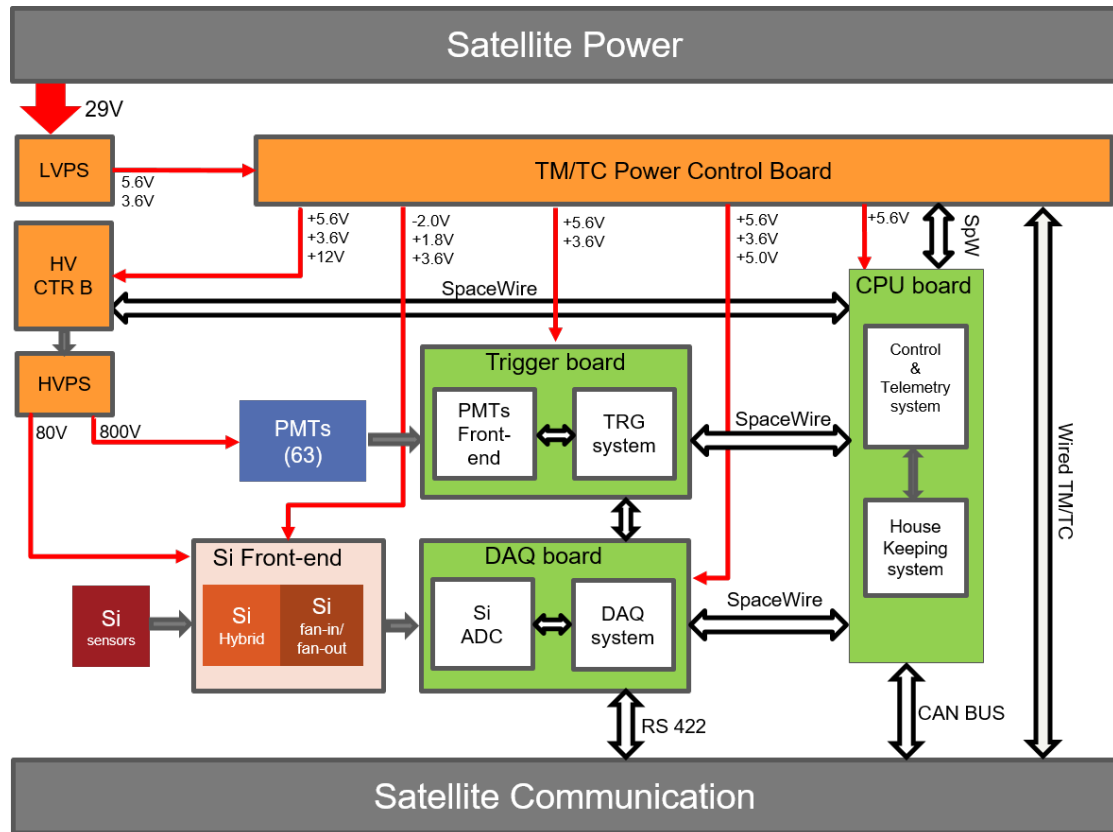


Figure 3: HEPD electronics block diagram - Connections with satellite and voltage distributions are shown.

- 195 • DAQ board - it hosts the Data acquisition system and the ADCs for the Silicon Front-end electronics; 218
- 196
- 197 • PMT/TRIGGER board - it hosts the front-end electronics of the PMTs and Trigger system; 220
- 198
- 199 • TeleMetry & TeleCommand (TM/TC) Power Control board (PCB) - it hosts a direct link to the satellite to receive specific hardware commands, and it distributes the digital voltages to all other electronic systems; it is part of the Power system; 223
- 200
- 201
- 202
- 203
- 204 • LVPS (Low-Voltage Power Supply) - it hosts the DC/DC converters (29 V to 5.6 V and 3.6 V, respectively) and provides the input voltages to the TM/TC PCB; it is part of the Power system; 228
- 205
- 206
- 207
- 208 • HV Control Board - it hosts the system controlling the High-Voltage Power Supply; it is part of the Power system; 232
- 209
- 210
- 211 • HVPS (High-Voltage Power Supply) - it hosts 10 HV units (0-1200 V) that provide bias voltages to the PMTs, and two HV units (0-150 V) that give voltage to the silicon sensors; it is part of the Power system. 236
- 212
- 213
- 214

215 Heat dissipation is managed by pure conduction through
 216 the aluminum mechanical modules where electronic boards
 217 are fixed. Board interconnections are provided by means of

a dedicated harness made of Glenair Micro-D connectors¹ and Glenair Ambestrand braided EMI shield².

Each electronic board hosts two identical copies of the same electronics (main/hot and spare/cold side) for redundancy. Main and spare sides are completely independent of each other and cannot be powered at the same time. A second level of redundancy has also been applied to some important components of each board.

All the programmable logic arrays (FPGAs) used in HEPD belong to the ProASIC3E Microsemi family³, which features several advantages for use in Space. The ProASIC3E family implements an on-chip non-volatile flash EEPROM that stores the configuration of the FPGA logic structure; this technology features good performance for what concerns radiation effects. For the programmable logic cells, there is no susceptibility to Total Ionization Dose (TID) up to at least 20 krad, while the overall cross-section for Single Event Effects (SEE) is of the order of 10^{-4} cm²/kbit, with a component of Single Event Latchup per device (SEL/device) that returns a negligible contribution to the overall SEE rate.

¹<https://www.glenair.com/micro-d/index.htm>

²<https://www.glenair.com/composite/d.htm>

³<https://www.microsemi.com/product-directory/fpgas/1690-proasic3>

239 On the other hand, the flash configuration system²⁷²
 240 (i.e., the part of the FPGA logic that is responsible for²⁷³
 241 overwriting the previously stored logic configuration with²⁷⁴
 242 a new one) has higher sensitivity to ionizing radiation,²⁷⁵
 243 but it is disabled and never used in flight.

244
 245 For the storage of non-volatile data (firmware as well as
 246 run-time data), the HEPD electronics relies on Ferroelec-²⁷⁷
 247 tric Random Access Memories (FRAMs) and flash type²⁷⁸
 248 memories. These FRAMs use a ferroelectric layer instead²⁷⁹
 249 of a dielectric one, nonetheless offering the same func-
 250 tionalities as flash memories, but with noticeable advan-
 251 tages over them. FRAM memories require lower voltage
 252 to perform a write operation, which implies lower power²⁸⁰
 253 consumption (1:21 ratio). Moreover, write operations are²⁸¹
 254 faster (16:1 ratio), and several tests show that *Single Event*
 255 *Upset*⁴ response is good. Another, and probably most sig-²⁸⁴
 256 nificant feature, is the device endurance: the number of
 257 write cycles is much higher than in flash memories and
 258 EEPROMs. The endurance to the write cycles is a crucial²⁸⁶
 259 issue in the HEPD framework, since the FRAMs are not²⁸⁷
 260 only used to store the program codes of the Digital Signal
 261 Processors (DSPs), but also detector calibration data that
 262 are calculated at every satellite orbit.

263 The following sections describe HEPD electronic sub-²⁹¹
 264 systems in detail.

265 4.2. Control system

266 The Control system is responsible for the management²⁹⁵
 267 of the detector and communications with the satellite plat-²⁹⁶
 268 form. It manages the following functionalities:

- 269 • Communication with Satellite OBDH computer via²⁹⁸
 270 CAN-bus (both nominal and redundant);²⁹⁹
- 271 – Management of TeleCommands (TCs);³⁰⁰

- Management of satellite information (broad-
cast);
- Management of TeleMetry data (Fast and Slow
TM);
- Storage of non-volatile information;
- Management of system configuration and in-flight op-
erations (orbital configurations, calibration and acqui-
sition runs);
- Configuration, test, and upgrade of the DAQ software
via the Mail box interface;
- Temporal tag of runs and broadcast (local time, ab-
solute time);
- Management, via SpaceWire link, of:
 - DAQ board;
 - PMT/TRIGGER board;
 - Low-Voltage control board (LVCB);
 - High-Voltage control board (HVCB).

289 The Control system is located in the CPU board. The
 290 board is divided into three areas: common, hot and cold.
 291 In the common area, the signals of the nominal and redun-
 292 dant CAN-bus links in the hot and cold areas are combined
 293 together by means of an impedance matching network.
 294 The electronics accommodated in the hot and cold areas
 295 are identical, and their architecture is based on the use of a
 296 Microsemi FPGA and a DSP of the class ADSP2189M by
 297 Analog Device⁵. The interconnection between the FPGA,
 298 DSP and various peripherals is shown in the block diagram
 299 in figure 5.

300 The main program that performs the board routines
 301 runs on the DSP, while the main functionalities of the
 302 FPGA are:

- 303 • Safe boot management;
- 304 • DSP Watch Dog management;
- 305 • Non-Volatile (NV) memory management;
- 306 • DATA memory (SRAM) management;
- 307 • SpaceWire Slow Control links management (see §4.2.1
 308 for details);
- 309 • CAN-bus links management (see §4.2.1 for details).

310 After boot, the Control system is responsible for
 311 the power-on sequence and initialization of HEPD sub-
 312 systems, following a proper sequence. The first step is

⁴A Single Event Upset (SEU) is a change of state in electronic
 devices caused by one single ionizing particle striking the component.³⁰³

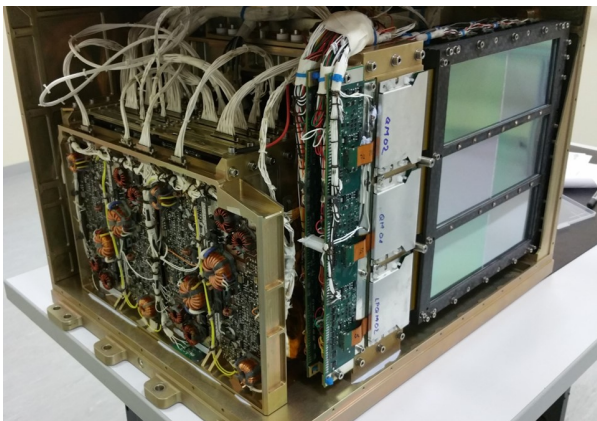


Figure 4: The electronics block and front-end electronics of the silicon tracker.

⁵<https://www.analog.com/en/products/adsp-2189m.html>

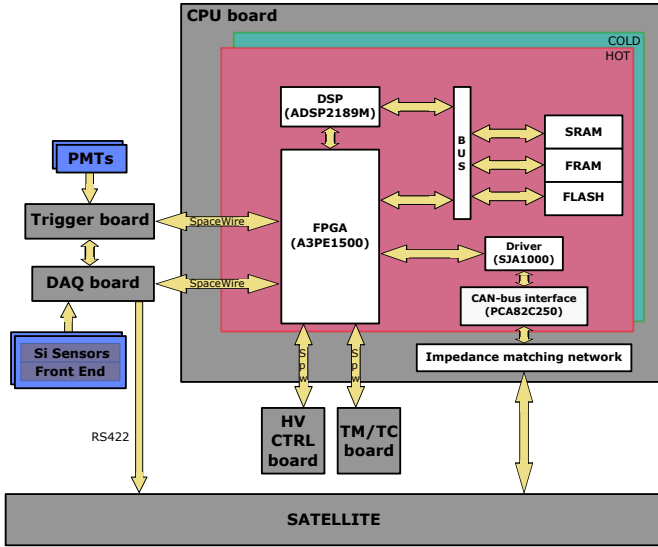


Figure 5: The CPU board block diagram

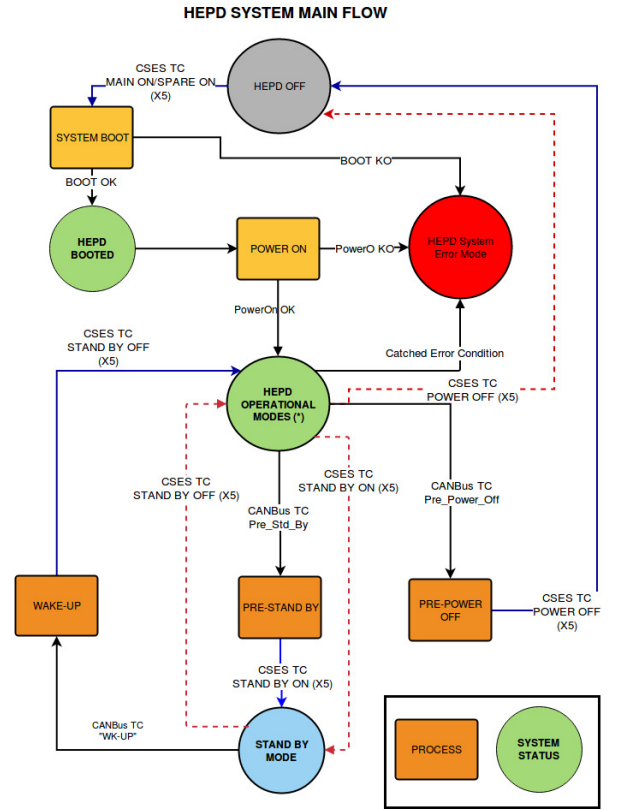


Figure 6: HEPD Main flow: any main HEPD status is shown, together with TCs required to status change.

313 to power each board on and immediately check its basic
314 functionalities.

315 If all boards are booted correctly, HEPD is set either in
316 NOMINAL or in SAFE mode (together called OPERA-
317 TIONAL modes) depending on the saved configuration; if
318 set in NOMINAL mode, the High-Voltage Control boards
319 are powered at their nominal values, providing bias to the
320 PMTs and Silicon ladders. At the same time, the Control
321 system stores non-volatile information, such as the Boot
322 Number and HEPD status before and after the first step,
323 into the FRAM. On the other hand, if set in SAFE mode,
324 the Low Voltage Power Supply and all boards are powered,
325 but the PMTs and Silicon planes are not biased; this
326 mode is mainly used for tests, in-flight health-check pro-
327 cedures at first power-on, and during some stages of the
328 commissioning phase.

329 Once in either OPERATIONAL mode, the Control system
330 starts a main loop during which data acquisition and
331 calibration runs are configured and automatically executed,
332 depending on the orbital zone configurations⁶.

333 From an OPERATIONAL mode, the HEPD can be
334 powered off or set in STAND-BY mode (used for non-
335 acquisition zones such as geographic poles, where plat-
336 form adjustments are performed) by means of dedicated
337 telecommands, received from the satellite OBDH.

338 Figure 6 illustrates the flow chart of any HEPD status,
339 the paths from any status to the others, and the corre-
340 sponding telecommands that trigger HEPD status change.

341 Four types of broadcast messages are periodically sent
342 by the satellite to all payloads on board CSES, providing
343 information about position, velocity, time and attitude:

⁶In SAFE mode, since HVPS are switched off, the Trigger board
generates fake triggers to test the whole electronic chain, including
data transmission to the satellite via the RS422 link.

344 OBDH time, GPS position and velocity, attitude and
345 star tracker data. Satellite latitude and longitude are
346 used by the Control system to identify the orbital region
347 of the satellite, to decide the next operation to execute
348 (either an acquisition or calibration run), and to change
349 the configuration of the apparatus. The satellite timing
350 information, together with CPU timestamp, is forwarded
351 to the DAQ system for the temporal tag of acquisition
352 and calibration runs.

353 During operations, the Control system updates the
354 HEPD configuration and receives information about the
355 status of each subsystem collected by the Housekeeping
356 system. The status of the apparatus is periodically
357 sent to the satellite as Fast Telemetry (Fast TM) and
358 Slow Telemetry (Slow TM) with a cycle of 1 s and 8 s,
359 respectively; the former basically contains the monitored
360 status register of each subsystem, while the latter includes
361 a larger set of parameters. All the acquired telemetries
362 are transmitted to ground once a day.

363 HEPD configuration and board status, as well as complete
364 broadcast and timing information, are stored into the
365 SRAM and sent to the DAQ via Mail Box to be integrated
366 into the scientific data.

367 The design and implementation of the software respon-
368 sible for HEPD management - i.e., handling instrument
369

371 data acquisition, performing periodic calibration of sub-416
372 detectors, monitoring system status, performing data com-417
373 pression, and communicating with the satellite - has been418
374 reported in detail in [7].

375 4.2.1. CPU interfaces 421

376 This section describes the different communication in-422
377 terfaces implemented on the CPU board for data exchange423
378 with the satellite (CAN-bus link), electronic boards (slow424
379 control link) and DSP of the DAQ board.

380 • CAN-bus interface 425

381 The physics layer of the interface has been designed in426
382 accordance with the CAN-bus 2.0 protocol. On the CSES429
383 satellite, two physical CAN-buses (A and B) are used for430
384 redundancy purposes; nevertheless, only one bus must be431
385 active at any time. In order to manage the CAN-bus,432
386 the registers of the two controllers (A and B) are memory433
387 mapped into the FPGA memory, and the DSP uses the434
388 I/O space interface to access and configure these registers.435

389 A custom driver has been designed, implemented and436
390 qualified to manage the two SJA1000 CAN-bus controllers437
391 by PHILIPS⁷ on the CPU board, considering the timing438
392 constraints defined for CSES by the satellite designers.439
393 The transceivers used for the CAN-bus electrical interface440
394 belong to the PCA82C250 class produced by PHILIPS⁸.441
395 The driver performs the following tasks:

- 396 • initialization and configuration of both controllers:443
397 i.e., setting the filter and bit rate according to CSES444
398 specifications;
- 399 • management of both CAN-bus channels;
- 400 • management of the HEPD CAN-bus protocol:
 - 401 – reception of four different types of Broadcast450
402 messages;
 - 403 – reception of Telemetry pool process messages451
404 and transmission of the Telemetry responses452
405 within prefixed time constraints (Fast/Slow453
406 TM);
 - 407 – reception of single- and multi-frame telecom-456
408 mand messages and transmission of telecom-457
409 mand message acknowledge within prefixed time458
410 constraints.

411 All the payloads on board CSES-01 can be considered460
412 as nodes on the CAN network. Each message has a well461
413 defined identification, priority and structure in order to be462
414 recognized by the payloads.

415 • Slow Control interface 463

The communication between the CPU and the electronic419
boards is based on the SpaceWire Light standard, accord-420
ing to which the controllers are implemented in the FPGA421
of each electronic board. Network nodes are connected422
through a serial link with low latency and allowed speeds423
between 2 and 200 Mbits/s. Four channels are managed:424
DAQ board, HV board, Power Control board and Trigger425
board.

The CPU board always acts as the master of the control426
link, while the other electronic boards act as slaves;427
each board is designed to export a memory mapped428
interface, which can be accessed by the CPU by means429
of the I/O memory interface. The slow control link is430
continuously used to read the status and error register,431
and to configure each board. Each link is managed by a fi-432
nite state machine as specified by the SpaceWire standard.

• Mail box interface 436

The last communication protocol (Mail box) regards437
data exchange between the DSP on the CPU board and438
the one of the DAQ board, the CPU one being the master439
of the protocol.

The protocol is implemented on the main FPGA of the440
DAQ board using 32-bit data registers and two control441
bits. The CPU accesses these registers via the slow control442
link, while the DSP of the DAQ uses the memory mapped443
space by means of the I/O interface. Specific registers are444
used to notify when a message from the CPU/DAQ must445
or can be read (i.e., the mail box is not busy), and to446
store data from the DAQ to the CPU or from the CPU to447
the DAQ. In order to avoid a conflict accessing the same448
register at the same time, the data register access from449
the CPU is possible after verification of the control bits,450
ensuring the mail box is not busy; once data are written,451
an interrupt is generated from the FPGA to the DAQ-DSP452
that extracts the data content, processes the command and453
writes a response onto a dedicated register.

454 4.3. Silicon Tracker Front-end and DAQ electronics 455

The readout of the Silicon Tracker is based on the456
IDE1140⁹ front end chip, a 64 channel low-noise/low-457
power high-dynamic-range charge-sensitive preamplifier-458
shaper circuit. The 64 channels are connected to an out-459
put buffer by means of an analog multiplexer that allows460
the sequential readout. A total of 36 chips are used to461
read out the two sides of the silicon sensors, 18 chips each462
side (ohmic and junction side), mounted on three hybrid463
boards. Due to the high density of the detector chan-464
nels, there is no redundancy. The six chips on each hybrid465
board are arranged in two groups for signal amplification466
and A/D conversion, as shown in Figure 7. Each hybrid467
board is connected to the DAQ board via a fan-in/fan-out
where decoupling circuits are presents for all the control

⁷<https://www.nxp.com/docs/en/data-sheet/SJA1000.pdf>

⁸<https://www.nxp.com/docs/en/data-sheet/PCA82C250.pdf>

⁹<https://ideas.no/products/ide1140/>

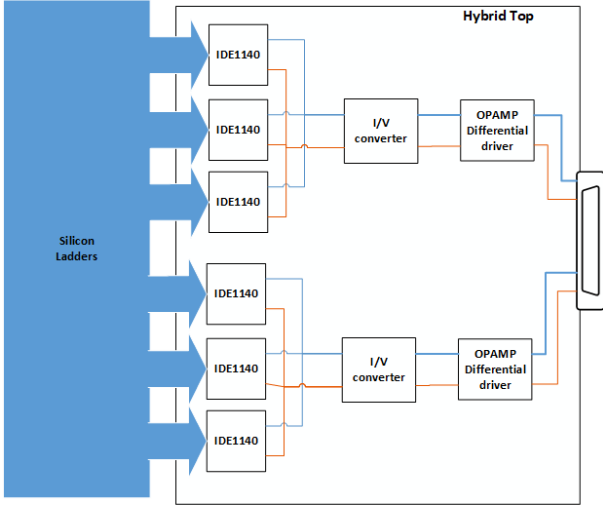


Figure 7: Hybrid Electronic block diagram. Only the analog signals are reported; the IDE1140 control signals are not shown.

signals generated by the DAQ board itself. The fan-in/fan-out board plays a crucial role also in power distribution: hot/cold redundant lines from the HVPS that bias the silicon detector, as well as hot/cold redundant lines from the LVPS that power the IDE1140 and the analog amplifier on the hybrid board, are all ORed in this board to properly power the hybrids and silicon detectors. Furthermore, this board receives the enable signals that allow to independently power on/off each detector column (two corresponding hybrids on both planes). Once a trigger is received, the DAQ board produces the HOLD signal for the IDE1140 chips, as well as the clock signal to control the analog output multiplexer. The amplified analog signals from the detector are digitized on the DAQ board using one ADC AD7274 every three front-end chips and producing a total of 27 kbit of raw data. The DAQ board applies a zero suppression algorithm in order to reduce data volume for transmission to ground.

4.4. Trigger/Calo/Veto front-end and DAQ electronics

The readout of the scintillator detectors (Trigger, Calorimeter and Veto) is performed through a dedicated board (Trigger Board) divided into two identical sections (Hot/Cold). Each section features an FPGA and two ASIC integrated circuits, followed by four ADCs. For redundancy, the signals coming from the PMTs placed on the same scintillator plane/paddle are driven to different front-end ASICs, with the exception of the LYSO plane where each cube is read out by only one PMT. In this way, a possible problem occurring in one of the ASIC chips does not harm the operation of the whole apparatus.

Figure 8 shows the block scheme of the Trigger Board. The board can handle 64 analog signals coming from the PMTs. The incoming signals are routed to a conditioning stage and then to two ASIC chips, which are used to integrate, shape and store the incoming signals. The analog signals coming from the ASICs are sent to the ADCs.

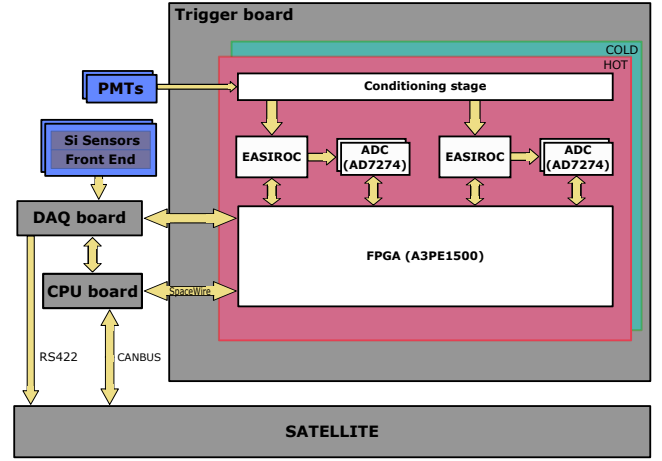


Figure 8: Block diagram of the trigger board.

The digitized signals are finally routed to the FPGA, which controls the whole process and handles the communication with the rest of the apparatus [8].

The ASIC performing the readout is the EASIROC by Weeroc¹⁰. This chip has been chosen because it represents a compact low-power solution to read out 32 channels (power consumption = 7 mW/channel). In addition, it has a wide dynamic range thanks to two independently-programmable variable-gain analog outputs (high gain and low gain), which offer multiplexed charge measurement from 160 fC up to 320 pC. These charge paths are composed of two variable gain preamplifiers, followed by two tunable shapers and a track-and-hold system that can be controlled by external signals. Slow shaping time can be adjusted between 25 and 175 ns for both low and high gain. A block scheme of the EASIROC chip is shown in figure 9.

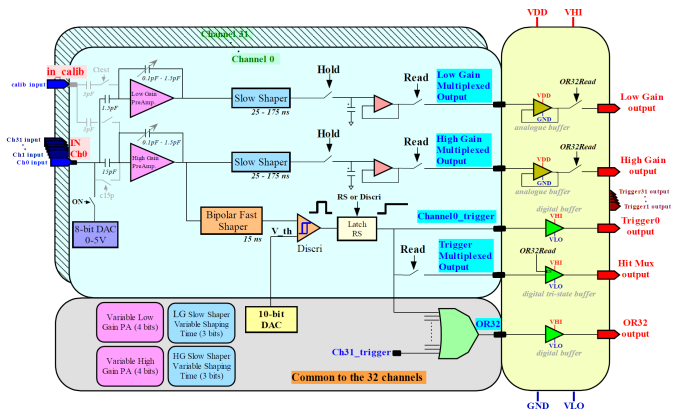


Figure 9: Block diagram of the EASIROC chip [9]

Since the EASIROC was originally developed to read out signals from silicon photomultipliers (SiPMs), there-

¹⁰<https://www.weeroc.com/>

fore requiring positive signals as an input, we have designed a conditioning stage to invert and attenuate analog signals from any PMT. The input attenuation is necessary because the signals produced by the photomultipliers can reach 5 V, which is well above the maximum allowed input signal to the ASIC. The implemented attenuation factor is 30.

The EASIROC chip is also equipped with a fast shaper, fed by the signals from high gain preamplifiers, in order to generate a fast trigger output for each channel with adjustable threshold. These outputs are directly routed to the FPGA, where they are used to drive a more complex trigger logic generation described in section 4.5. The generated trigger signal drives data acquisition of the signals from the PMTs, allowing the track-and-hold cells of the EASIROC to save the amplitude of the preamplified and shaped signal at its peaking time.

The 32 high and low gain channels are then driven to two different ADCs to convert the EASIROC readout signals into digital ones. The ADCs used on the Trigger Board are produced by Analog Devices, Inc., model AD7274¹¹. These ADCs are 12-bit, high-speed, low-power, successive approximation ADCs, and they feature throughput rates of up to 3 MSPS.

For each event, the digitized signals from the PMTs (12×32 = 384 bits) are associated with other scientific data, such as the rate meters for each trigger configuration (8×32 bit counters), and the dead and live times of the whole instrument. These data are transmitted to the DAQ Board following the protocol described in section 4.6.

A slow control and command interface with the CPU has been designed and implemented to configure the EASIROC, and to configure the trigger generation algorithm. The acquisition and calibration runs are also controlled through this interface, described in section 4.2.1.

Finally, the Trigger Board implements 63 additional counters to measure single PMT “rate meters”, which are transmitted along with scientific data. These counters measure the trigger rate of every PMT over 1 second, and allow to understand whether a single PMT is damaged or misbehaving, such that it can be masked and ignored in the generation of the trigger configuration.

4.5. Trigger system

The Trigger Board implements the trigger system of the apparatus. Besides the analog output, the EASIROC chip provides 32 individual digital output signals, one for each channel, and an additional signal that is generated via the logical OR of the single channel triggers (“OR32” in figure 9). These outputs are produced every time a signal exceeds a certain threshold value on the corresponding input. The threshold value is the same for all the channels, it can be changed through a CPU command, also from ground, by sending a command to the Trigger board through the slow

control. The FPGA firmware is configured to issue a global trigger each time the trigger pattern of an event complies with a mask that can be chosen in a predefined set. When a global trigger is generated, the Trigger board starts the handshaking process described in section 4.6, in order to start the acquisition of scientific data from the tracker and the scintillator detectors.

HEPD can tap into 8 predefined “trigger masks” plus a configurable Generic Trigger Mask. Said T the “OR” of the six T_i trigger counters, P_i the i_{th} scintillator plane of the calorimeter, and L the “OR” of the 9 LYSO crystals, the eight masks are obtained by different logic combinations of T, P_i and L. The Generic Trigger Mask can be configured as any “AND” combination of calorimeter planes and trigger counters.

The different trigger masks define the aperture and the energy acceptance of the instrument: by requiring a deeper penetration of the particle inside the detector (i.e., using the trigger counters and a larger set of P_i scintillator planes in “AND” configuration), the geometric factor of HEPD decreases and the energy threshold for triggering increases.

Each of these predefined trigger masks can be used with different VETO settings: no veto, lateral veto alone, bottom veto alone, whole veto (lateral+bottom).

For each of the predefined masks, even when not selected for the online acquisition, a rate meter provides the corresponding trigger counting rate, allowing to simultaneously give independent estimates of count rates of different particle populations crossing the instrument, indirectly limiting the energy thresholds. These data are provided for each event and are part of the scientific data.

Depending on the selected HEPD acquisition mode, the Trigger system can work in different modes: either “Calibration” or “Event Acquisition” mode.

In Calibration mode, the Trigger board generates and sends fake trigger signals to the DAQ board, which acquires the ADC signals from the scintillators and silicon detector. Such data are sent to the DSP of the DAQ board for processing, in order to evaluate the pedestal, RMS and status of each silicon detector strip and each PMT. These calibration data are used online for the silicon detector data reduction, and offline for data analysis.

In Event Acquisition mode, the Trigger board generates and sends triggers to the DAQ board according to the specified threshold of the PMTs, on-line trigger mask configuration and VETO setting. The PMT signals, digitized in the Trigger Board, are then sent to the DAQ, where silicon data are acquired and processed.

In Event Acquisition mode, the Trigger Board also activates two counters to measure the dead and live times of the apparatus, which are appended to the scientific data sent to the DAQ. The calculation of both dead and live time of the apparatus is fundamental to evaluate the fluxes of incoming particles.

¹¹<https://www.analog.com/en/products/ad7274.html>

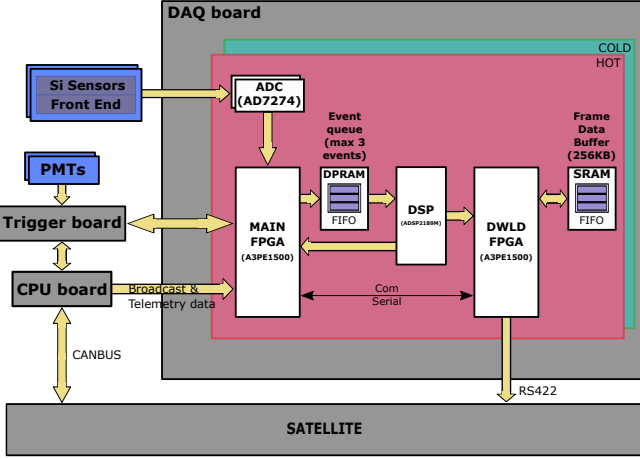


Figure 10: The Data Acquisition System involves DAQ board, Trigger board, CPU board and front-end electronics of the silicon and PMT detectors. All the scientific data are transferred to the DAQ board, where they are processed and compressed before transmission to the satellite via the RS422 link.

4.6. Data Acquisition System

The Data Acquisition System is responsible for the acquisition and management of the scientific data, as well as the processing of the digital signals and their transmission to the satellite. It is located in the DAQ board, but it also involves the CPU and Trigger board. The DAQ system manages the following main functionalities:

- configuration of the tracker detector for acquisition runs (Event Acquisition or Calibration mode);
- handshake with the Trigger board for the generation of the trigger pulse, coordination of all sub-systems for event acquisition, and calculation of the dead/live time of the apparatus;
- acquisition of tracker data and consequent signal compression;
- acquisition of PMT data from the Trigger board;
- scientific data compression and formatting;
- temporal tag of the event;
- transmission of the compressed scientific data to the satellite.

A block diagram of scientific data handling and all sub-systems responsible for acquisition is shown in Figure 10. The DAQ board is the main element: it contains a section with all the analog and digital circuits for the interface with the silicon front-end and the digitization of the silicon data; two FPGAs (called MAIN and DOWNLOAD); and a DSP. The other components of the DAQ are: two FRAMs used for the storage of the code and other ancillary data (calibrations) that must

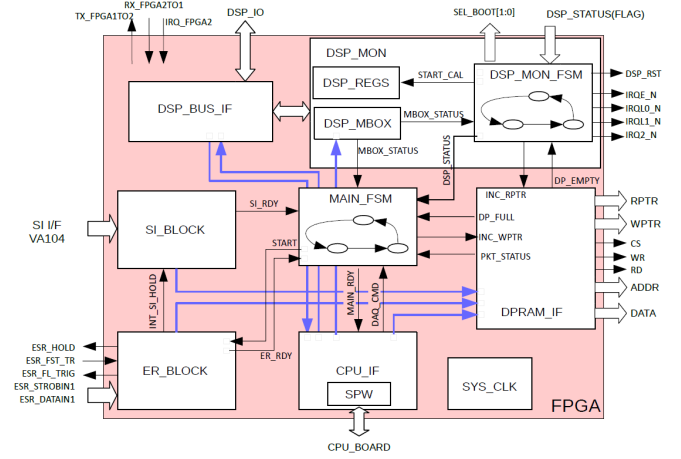


Figure 11: Block diagram of the firmware implemented on the Main FPGA. It is divided into six macro blocks, each one implementing a specific task: *SI_BLOCK* controls the silicon data acquisition and digitization; *ER_BLOCK* handles the communication with the Trigger Board for trigger generation; *CPU_IF* handles the communication with the CPU board; *DPRAM_IF* regulates the access to the DPRAM for the FPGA and DSP; *MAIN_FSM* is the finite state machine of the FPGA; *DSP_MON_FSM* is the monitor of the DSP finite state machine.

be preserved after board power-off; a Dual-Port Random Access Memory (DPRAM) used for data exchange between the MAIN FPGA and the DSP; a static RAM (SRAM) used as a FIFO buffer for the final output of HEPD scientific data that must be transferred to the satellite. The operating frequency of the board is 48 MHz, good enough for computing power needs while keeping an acceptable electric power consumption.

When HEPD operates in NOMINAL mode, the generation of a trigger pulse is coordinated by the DAQ and Trigger systems through a handshake procedure. If a particle releases an above-threshold signal on the PMTs involved in the trigger mask (as defined in 4.5), the Trigger Board asserts the *fast_trigger_signal*, to which the DAQ responds with an acknowledge and inhibits the generation of new triggers. The acquisition of PMT and silicon signals starts. The PMT signals are digitized directly on the Trigger board, and then sent to the DAQ by using a simplified version of the serial protocol adopted for the CPU link. The acquisition of the silicon detectors is coordinated by a dedicated firmware block on the MAIN FPGA of the DAQ (*SI_BLOCK* in Figure 11).

The readout and following digitization of the Silicon tracker data are the main source of instrumental dead time. Other contributions to dead time come from the transfer of digitized PMT data from the Trigger Board to the internal memory of the MAIN FPGA, the write operation of silicon and PMT data to the DPRAM, and event compression performed by the DSP itself.

It has been observed that data transmission from the Trigger Board via the serial link induces a noise on the

691 readout of the silicon signals, resulting in a significant⁷⁴⁶
 692 worsening of the RMS values of each channel. For this⁷⁴⁷
 693 reason, the two operations are serialized, in spite of a⁷⁴⁸
 694 slight increase in the total dead time. The time neces-⁷⁴⁹
 695 sary to acquire and write all the detector data into the⁷⁵⁰
 696 DPRAM, ready to be read out and processed by the DSP,⁷⁵¹
 697 is fixed because it only depends on the clock of the board.⁷⁵²
 698 This time has been found to be $\simeq 2.5$ ms. The access to⁷⁵³
 699 the DPRAM by the MAIN FPGA and DSP is managed
 700 and synchronizes through a handler present in the FPGA⁷⁵⁴
 701 (*DPRAM_IF* in Figure 11).⁷⁵⁵

702 This DPRAM works as a buffer that can contain up to⁷⁵⁶
 703 8 events. When the FPGA copies the acquired data to the⁷⁵⁷
 704 DPRAM, it becomes also free to process a new trigger in⁷⁵⁸
 705 parallel with the DSP processing the previous one. Only⁷⁵⁹
 706 in the case when all the 8 DPRAM pages are full, the DSP⁷⁶⁰
 707 stops the acquisition of a new event. The DPRAM acts⁷⁶¹
 708 like a FIFO where the first event written by the FPGA will⁷⁶²
 709 be the first event processed by the DSP. Another firmware⁷⁶³
 710 block in the MAIN FPGA (*DPS_MON* in Figure 11) op-⁷⁶⁴
 711 erates as a monitor of the DSP; it controls the finite state⁷⁶⁵
 712 machine of the processor by sending interrupt signals or⁷⁶⁶
 713 resetting the DSP in case of an unexpected change of state.⁷⁶⁷

714 The DSP compresses the Silicon data (reduced by a fac-
 715 tor $\simeq 80$) using a custom “zero suppression” algorithm.⁷⁶⁸
 716 Processed events are continuously stored into the SRAM;
 717 as soon as 4110 bytes are written, the DOWNLOAD⁷⁶⁹
 718 FPGA sends a packet to the satellite via the RS422 link,⁷⁷⁰
 719 adding a frame header and a second checksum at the end⁷⁷¹
 720 of the frame. In this way, event processing and transfer to⁷⁷²
 721 the satellite are carried out in parallel, and all RS422 pack-⁷⁷³
 722 ets have the same fixed dimension as required by satellite⁷⁷⁴
 723 specifications. The end of an acquisition run is handled by
 724 the CPU, which stops the acquisition after a given time⁷⁷⁵
 725 window or when specific orbital conditions are met (e.g.,
 726 when HEPD is about to be set in STAND-BY mode for⁷⁷⁶
 727 satellite attitude adjustment operations).⁷⁷⁷

728 4.7. Housekeeping system⁷⁷⁸

729 The Housekeeping system is responsible for HEPD diag-⁷⁷⁹
 730 nostic routines, and, like the Control System, it is imple-⁷⁸⁰
 731 mented as an application program running on the CPU.⁷⁸¹

732
 733 During the power-on and operations phases, the House-⁷⁸³
 734 keeping system periodically polls each electronic subsystem⁷⁸⁴
 735 via the slow control link, ensuring complete monitor-⁷⁸⁵
 736 ing of the whole detector.

737 The Housekeeping system also periodically collects infor-⁷⁸⁶
 738 mation about the High Voltage Power Supply modules - by⁷⁸⁷
 739 reading each HV value - and the temperatures recorded by⁷⁸⁸
 740 the sensors placed on the CPU and Trigger boards. Such⁷⁸⁹
 741 information is sent to the Control System and then relayed⁷⁹⁰
 742 to the satellite through Fast Telemetry (TM) and Slow TM⁷⁹¹
 743 with cycles of 1 s and 8 s, respectively.⁷⁹²

744 The Fast TM contains the monitored status register of
 745 each subsystem, while the Slow TM includes more param-

eters, such as error registers for each electronic board, the
 value of the temperature sensors placed on the CPU and
 Trigger Boards, the monitored values of the high voltages
 for the PMT and silicon detectors, and the last TC re-
 ceived.

Part of the monitored data (the board status) is also sent
 to the DAQ via the Slow Control link to be integrated into
 the scientific data (Run Header and Tail).

At the same time the Housekeeping system monitors the
 good functionality of PMTs by reading the 63 single PMT
 rate meters provided by the Trigger board. Those values
 can be analyzed offline to identify broken or bad working
 PMTs (for example, noisy PMTs) that could affect the
 Trigger system efficiency; these PMTs can be excluded
 from the trigger by means of a dedicated configuration
 telecommand.

The collected PMT information, as well as temperature
 and timing, are sent to the DAQ via the Slow Control
 link in order to be integrated into the scientific data.

A block diagram of the housekeeping data handling and
 transmission is shown in Figure 12.

4.8. Power system

The power supply subsystem provides low voltages
 (main power supply unit, LVPS) to the detector electron-
 ics, and the high voltage (HV) bias for PMTs and silicon
 planes (secondary power supply unit, HVPS). A schematic
 diagram of power supply distribution is shown in Figure
 3.

4.8.1. Low-Voltage Power Supply (LVPS)

The Low Voltage Power supply (LVPS) unit includes:

- two LV modules, customized versions of CAEN¹²
 S9074 and S9053 DC/DC converters, which receive
 the main 29 V power line input from satellite and
 deliver the appropriate voltages (5.6 V and 3.6 V,
 respectively) to the rest of the system via the Low
 Voltage Control Board (LVCB);
- the LVCB, which a) communicates with the satellite
 for wired telecommands and telemetries concerning
 the status of the LV modules, and b) generates and
 controls the various supply voltages for the rest of the
 system.

For the sake of redundancy, each LV module is doubled,
 while the LVCB consists of two identical parts (hot and
 cold).

The LV module is a high-reliability DC/DC converter
 designed for space applications; the output current is

¹²<https://www.caen.it/>

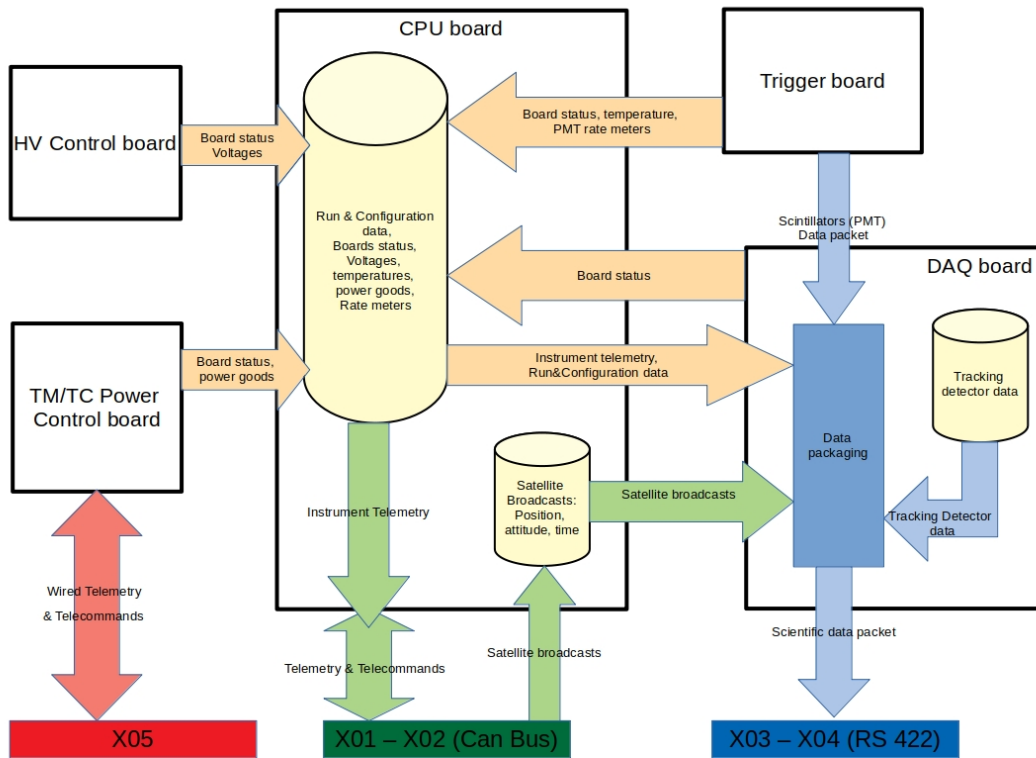


Figure 12: Block diagram of scientific and housekeeping data handling and transmission.

793 pulse-by-pulse controlled, in order to protect the circuits⁸¹⁴
 794 from overload and short circuit. The main features of the
 795 module are summarized here:

- 796 • radiation tolerance up to 30 krad;
- 797 • device protection to keep operating parameters within⁸¹⁸
 798 safe limits: input over/undervoltage control, out-⁸¹⁹
 799 put overcurrent control, maximum duty-cycle control,⁸²⁰
 800 solid-state fuse;
- 801 • guaranteed DC isolation between power input and⁸²²
 802 output circuitry up to voltage differences of 250 V;
- 803 • output voltage temperature stability within 0.5⁸²⁵
 804 mV/°C;
- 805 • line regulation better than 35 mV and 2 mV for the 5.6⁸²⁸
 806 V and 3.6 V module, respectively, with input voltage⁸²⁹
 807 in the range from 26.5 V to 30.5 V;
- 808 • load regulation better than 220 mV and 130 mV, re-⁸³²
 809 spectively, over the full dynamic range and recom-⁸³³
 810 mended temperature range (-20°C to +70°C);
- 811 • peak-to-peak output ripple < 40 mV and 20 mV, re-⁸³⁵
 812 spectively, in the bandwidth up to 20 MHz;
- 813 • typical power efficiency of 79% at full output power.

4.8.2. High-Voltage Power Supply (HVPS)

815 The High Voltage Power Supply (HVPS) unit is com-
 816 posed by the following electronic subsystems assembled in
 817 a common metallic frame:

- 10 step-up HV modules with negative output (up to 1200 V DC) for PMTs (each module serves several PMTs between 4 and 7);
- 2 step-up HV modules with positive output (up to 150 V DC) for the silicon planes (one module serves one silicon plane);
- HV control board (HVCB) to set and monitor the HV output values and interface with the CPU board.

826 A block diagram of the HVPS is shown in Figure 13.

827 Like LV modules, each HV module is doubled for redun-
 828 dancy, while the HV control board comprises two identical
 829 parts (hot and cold). Each pair of hot/cold HV output ca-
 830 bles are short-circuited into a dedicated diode protection
 831 circuit PCB housed externally to the HVPS on the sup-
 832 port mechanics of the unit, which produces a single HV
 833 output line in turn split into a number of cables adequate
 834 to reach the corresponding loads.

835 The employed HV module is Aerospazio HV3, a high-
 836 reliability HV DC/DC converter¹³ unit designed for space

¹³https://www.aerospazio.com/high-voltage_DC-DC_converters.html

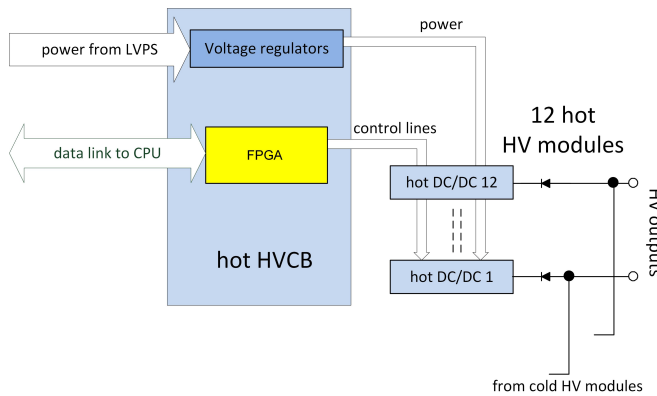


Figure 13: Block diagram of the HVPS hot part.

applications, optimized for low ripple, low power consumption, wide operating temperature range, compact size and light weight. The circuit design of this unit is suitable for generation of both negative and positive HVs, for PMTs and silicon planes, respectively. The main features of the module are summarized here:

- maximum HV output: ± 1200 V;
- output voltage analog monitor line;
- programmable output current limiter (maximum $500 \mu\text{A}$) with overcurrent status and current monitor lines;
- output load regulation stability within 0.4%;
- output voltage linearity better than 0.4%;
- output voltage temperature stability better than $100 \text{ ppm}/^\circ\text{C}$;
- peak-to-peak output ripple $< 25 \text{ mV}$ (in the bandwidth 20 Hz to 20 MHz).

To guarantee proper operation under any pressure condition from room to space environment, and taking into account the compactness of the unit which implies reduced distances between circuit elements, the HV module PCB has been installed in a housing frame filled with potting substance (Arathane 5753-A/B LV) that assures a high enough dielectric constant to avoid discharges between the circuit parts at different voltages. The low-outgassing properties of the potting material are compatible with the requirements for space application.

The main functionalities of the HVCB allow to:

- configure the HV value (with 0.1% resolution) for single HV modules by means of dedicated voltage control lines;
- switch on/off single HV modules by means of dedicated enable lines;

- check for SEUs in the critical voltage setting registers;
- interface with the CPU board (SpaceWire light link).

The setting of the HV voltage for a given channel is operated by the HVCB on-board logics either automatically at power-on or when a suitable command is sent by the CPU board. The change in the HV voltage can be operated through a sequence of steps of max 300 V at 1 s time intervals, to properly limit the inrush current at each step. The HV output resolution is 0.1%, while the combination of channel fabrication tolerances and non-linearities returns a total uncertainty $< 1\%$ on the actual HV value with respect to the nominal one.

The setting of the HV output is operated on the HVCB by driving the voltage set analog input V_SET of the HV module through the output of a RC-RC filter (time constant order of 1 ms). The RC-RC, in turn, is fed by a by an FPGA output producing a square wave with fixed period ($22 \mu\text{s}$, i.e., much smaller than the filter time constant) and variable duty cycle; the filter output is therefore a voltage level that is proportional to the chosen duty cycle. With this method, the accuracy of the analog V_SET value is given by the accuracy of the duty cycle, which has negligible dependence on environmental conditions and aging of devices.

Though SEUs in HVCB logics are highly improbable with the expected in-orbit charged-particle fluxes, each HV setting register is individually protected by an Error Detection And Correction (EDAC) Hamming function, with continuous (each 50 ns) automatic correction of single bit upset and detection of double bit upset (with internal alarm raised while the HV setting is automatically brought to 0). The fast correction rate is such that, taking into account the RC delays in the HV line, any upset has no significant effect on the corresponding HV output voltage.

The actual HV value of each channel can be monitored through both the readout of the setting register and the analog monitor output of the HV module. This monitor output voltage linearly replicates the HV output (in a voltage range scaled by a factor 300) and is in turn digitally converted by a dedicated 10-bit ADC (Analog AD7276¹⁴) to fill the AN_MON internal register of the FPGA. The overall combined fabrication tolerance + non-linearity of the monitor line with respect to the actual HV output voltage is within 3%.

At power-on, the on-board logic performs a fixed sequence of operations, in such a way that, even in the absence of any command from the CPU board (because of a failure of the command interface), each HV channel is set to an adequate default value: the default values are loaded during the final FPGA configuration (HVCB acceptance test). Though they differ from the optimized values, they are still capable of assuring satisfying operation

¹⁴<https://www.analog.com/en/products/ad7276.html>

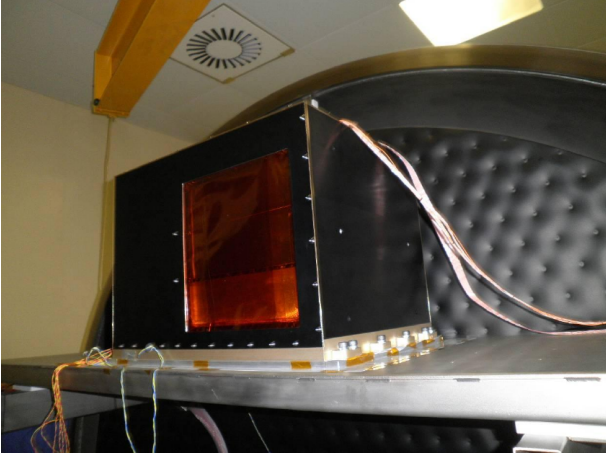


Figure 14: HEPD Flight Model (FM) in the thermal vacuum chamber at the SERMS facility in Terni (Italy) during the payload acceptance campaign in November 2016.

of the PMTs and silicon planes. The automatic sequence with ramp-up of HV outputs to default values can be interrupted if a suitable command from the CPU board arrives within a predefined time (4 s) from HVCB power-on.

The interface with CPU boards allows commands to be received to switch on/off single HV modules, to set HV values, and to read voltage monitors and alarm conditions (such as double-bit upset in the HV set register or failure in the board power circuitry).

5. Test and Qualification Campaigns

In compliance with the model philosophy of the project arranged with the Chinese DFH Satellite Co. - located in Beijing (China) - four HEPD models have been developed.

The Electrical Model (EM) is a mock-up of HEPD where all electric and electronic interfaces (power bus, TM/TC, CAN-bus and RS422) with satellite are developed. This model was realized to demonstrate the hardware and software design of the payload, to verify the electric and electronic compatibility between payload and satellite, and to test the compatibility between HEPD and its Electrical Ground Support Equipment (EGSE). The HEPD EM was successfully tested at the DFH test facility in October 2014.

The Structural and Thermal Model (STM) is a complete mechanical mock-up with dummy sensors and electronics in place of electronic boards to emulate the real heat dissipation. This model was developed to validate the structural and mechanical design, to test the payload thermal control design, and to verify the compatibility between HEPD and its Mechanical Ground Support Equipment (MGSE). Also the HEPD STM was successfully tested at the DFH test facility in October 2014.

The Qualification Model (QM), identical to the flight version of the payload, was developed and submitted to a

complete qualification test campaign to assess the design and the technological solutions, and to demonstrate its performance.

In compliance with the environmental test requirements, the QM underwent an extensive qualification test campaign, which included pyroshock, sinusoidal and random vibration, thermal-cycling and thermal vacuum tests. These tests were carried out at the SERMS Laboratory in Terni (Italy) between May and August 2016.

Specifically, in the pyrotechnic shock test HEPD QM was exposed to two different shocks along each of its three axes. The Shock Response Spectrum (SRS) of this test had frequencies ranging from 600 to 4000 Hz and acceleration up to 1000 g. A Sinusoidal vibration test was performed along the three axes with amplitude 7.5 mm in the frequency range 10-20 Hz and 12 g acceleration in the frequency range 20-100 Hz with a scan rate of 2 oct/min. A Random vibration test was carried out along the three axes in the frequency range 10-2000 Hz with a total GRMS 11.3 g and 2 minutes of loading duration.

The thermal-cycling test was performed in a climatic chamber at ambient pressure: 25.5 cycles were run in a temperature range between -30°C to $+50^{\circ}\text{C}$ with a temperature gradient of $(3-5)^{\circ}\text{C}/\text{min}$ and a dwell time of ≥ 4 hours. The dwell time is the minimum required time to achieve the thermal stabilization of the tested device.

The thermal vacuum test was carried out in a large vacuum chamber at a pressure $\leq 6.66 \times 10^{-3}$ Pa: six cycles were run in a temperature range between -30°C and $+50^{\circ}\text{C}$. The chamber temperature gradients were $2^{\circ}\text{C}/\text{min}$ and $1^{\circ}\text{C}/\text{min}$ during the heating and cooling, respectively, with a dwell time of ≥ 4 hours.

Each cycle of thermal-cycling and thermal vacuum tests included different steps corresponding to different operations for HEPD, such as data acquisition, calibration, stand-by or power off, in order to simulate in-flight procedures. During these operations, HEPD telemetry packets were continuously monitored to check for anomalies. The EGSE was used to power and manage HEPD, in order to change its status and operating mode according to test requirements.

Once the QM successfully passed the qualification test campaign in Italy, it was delivered to DFH. Then, HEPD QM was installed on the QM model of the satellite and successfully tested at the DFH test facility in October 2016. It was returned to the Italian side the following month.

The acceptance campaign for the Flight Model (FM) was carried out between October and November 2016. In compliance with ECSS (European Cooperation for Space Standardization) standards, no pyroshock test was performed. All other tests were performed under the same conditions as the QM, but with lower qualification levels or temperatures: a sinusoidal vibration test was performed along the three axes with amplitude 5.0 mm in the frequency range 10-20 Hz and 8 g acceleration in the

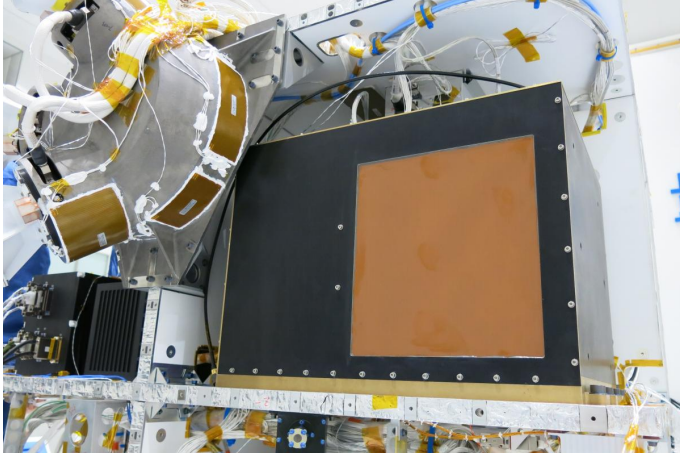


Figure 15: HEPD Flight Model installed on CSES satellite at DFH Satellite Company, Ltd. (Beijing, China)

1011 frequency range 20-100 Hz with a scan rate 4 oct/min.
 1012 A random vibration test was carried out along the three
 1013 axes in the frequency range 10-2000 Hz with a total GRMS
 1014 7.55 g and 1 minute of loading duration. For the thermal-
 1015 cycling test, 17.5 cycles were run in a temperature range
 1016 between -20°C to $+45^{\circ}\text{C}$. For the thermal vacuum test,
 1017 4.5 cycles were run in a temperature range between -20°C
 1018 and $+45^{\circ}\text{C}$. Figure 14 presents the HEPD Flight Model in
 1019 the thermal vacuum chamber at the SERMS facility in
 1020 Terni during the payload acceptance campaign in Novem-
 1021 ber 2016.

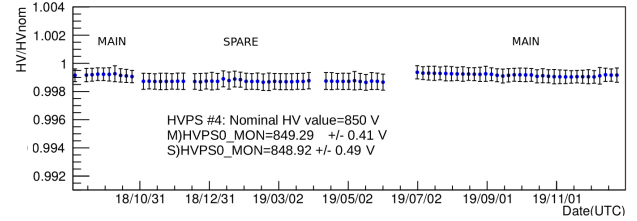
1022 The acceptance test campaign was successfully passed
 1023 by HEPD FM. Therefore, it was delivered to DFH in De-
 1024 cember 2016.

1025 In January 2017, HEPD FM successfully passed a stand-
 1026 alone assembly and integration verification procedure car-
 1027 ried out by means of its EGSE. Then, it was cleared for
 1028 the installation on CSES satellite at DFH, as shown in fig-
 1029 ure 15. Sinusoidal and random vibration tests, as well as
 1030 thermal-vacuum and thermal balance tests, were success-
 1031 fully performed on board CSES satellite between February
 1032 and April 2017. Magnetic cleanliness and aging tests were
 1033 completed in May 2017.

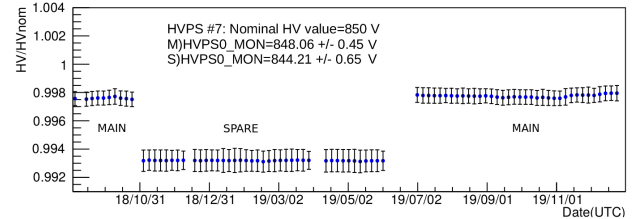
1034 Before delivery to China, the HEPD FM was tested un-
 1035 der particle beams in different beam test facilities. In Oc-
 1036 tober 2016, the detector was exposed to electron beams of
 1037 30, 60, 90 and 120 MeV at the INFN-LNF BTF (Beam
 1038 Test Facility) in Frascati (Italy). In November 2016, it
 1039 was exposed to proton beams of 37, 51, 70, 100, 125, 154,
 1040 202, and 228 MeV at the Trento Protontherapy Center in
 1041 Trento (Italy). Results from the test-beam calibrations
 1042 of HEPD can be found in [10].

1043 6. In-flight performance

1044 CSES satellite was launched from the Jiuquan Satellite
 1045 Launch Center in the Gobi desert (Inner Mongolia, China)



(a) HVPS 4



(b) HVPS 7

Figure 16: Variation of Monitored HV values with respect to the nominal ones for main/spare sides; mean values over the whole period for each side are also specified. Monitored HV values have been averaged over 5 days for the period September 2018- December 2019.

on February 2nd, 2018. Four days later, the HEPD instrument was switched on for the first time. All detectors were functioning properly.

After launch, a long commissioning phase (February–July 2018) started, during which several on-board configurations were implemented and tested in order to optimize the operational parameters of the instrument. Since August 2018, HEPD has been in science run and data-taking mode.

6.1. Housekeeping data

The status of the apparatus is continuously monitored by the Housekeeping System. The Telemetry data are mostly used to check the stability of the electronics, in order to identify anomalous behaviors and - if necessary - to investigate the possible source of malfunctions. An example of the continuous monitoring of the high voltage values that bias the PMTs is given in figure 16, where monitored HV values divided by the nominal values are shown over more than 1 year of life (September 2018 - December 2019), for two HVPS (no.4 and no.7) taken as an example. The HVPSs show a very stable behavior, with a variation below 0.5%. The discontinuity visible in the period November 2018 to July 2019 is due to the switch from the main to the spare side of the electronics, with different offsets for the two different High Voltage Power Supplies.

6.2. Scientific data

In addition to housekeeping data, also the scientific output is continuously monitored in order to perform offline calibration and check the stability of the detector response.

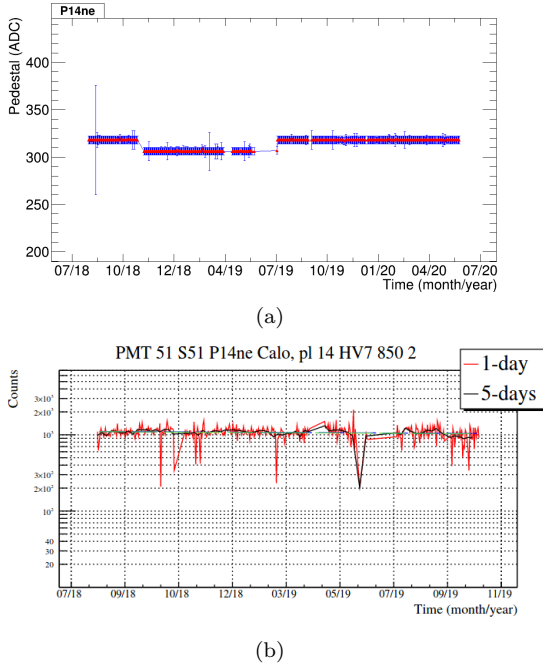


Figure 17: a) Pedestal of a specific PMT in plane 14 of the calorimeter, as calculated online during the calibration procedure. The steps at November 2018 and July 2019 are due to transitions from main side to spare electronics and viceversa. b) Rate meter for the same PMT by average over 1 day (red curve) and 5 days (black curve).

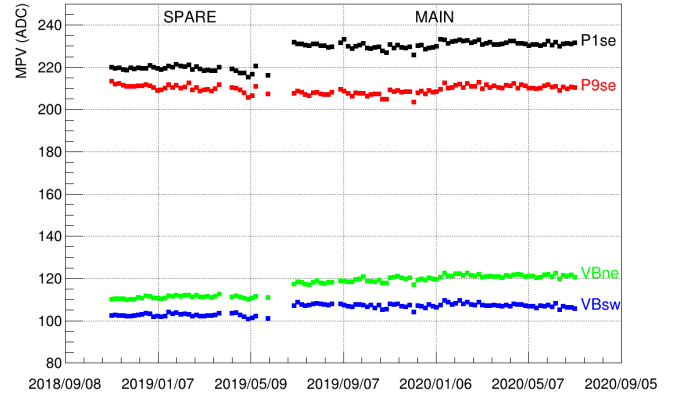


Figure 18: “mip” response as a function of time (from November 2019 until September 2020, in spare and main electronic acquisition) for four PMTs; one concerning the first calorimeter plane (black line), one for the 9th plane (red line), and two for the bottom veto plane (green and blue lines).

From the end of the commissioning operations to the end of June 2020, more than 10000 hours of acquisition have been collected, corresponding to about 10000 CSES orbits.

PMT output is checked by studying mean and variance of the pedestal for every channel. These quantities are calculated for each CSES orbit during the normal functioning of HEPD, and they show a very good stability over the entire period. As an example, in figure 17a, the pedestal of one of the two PMTs of the 14th plane is reported as a function of time. Each point corresponds to the value of the pedestal obtained from a single calibration run. A sudden decrease of the pedestal during the period from November 2018 to July 2019 can be observed, due to the switch from the main to the spare side of the electronics, while no significant variations are visible when the pedestals are measured under the same electronic conditions. Figure 17b shows the rate meter (i.e., the number of events that released a signal above the PMT threshold in 1 second) for the same PMT, by averaging over a single day (red curve) and five days (black curve). The single day trend shows some sharp depressions that can be explained either by intervals in which the detector was powered off or missing data (for example, because of data corruption during transmission to the ground). The black curve shows a rather constant trend, as expected for a PMT in the 14th plane, placed deep inside in the calorimeter, and for this reason not sensitive to fluctuations of low-energy particle fluxes (e.g., during geomagnetic storms or solar events).

The acquisition of particles with relativistic speed, called minimum ionizing particles (“mip”s), is also used to monitor and calibrate the PMT energy response. Although it is not possible to individually select mips by means of HEPD (due to the impossibility to measure the energy of particles not fully contained within the calorimeter), events passing through the entire instrument and bottom veto plane without releasing signal on the lateral veto, represent a sample statistically dominated by mips. Figure 18 shows the mip response as a function of time (from November 2019 until September 2020, in both spare and main electronic configuration) for four PMTs; one concerning the first calorimeter plane (black line), one for the 9th plane (red line), and two for the bottom veto plane (green and blue lines). On the Y axis the Most Probable Value (MPV) of the PMT ADC count is reported. The MPV is obtained by a Landau fit on the PMT global signal distribution. The MPV behavior looks stable over the period considered, suggesting a stable state of the electronics over time.

Eight rate meters, one for each trigger mask defined in §4.5, are also implemented. In figure 19, three of them are reported for the configurations:

- T (only trigger plane);
- T & P1 (trigger plane and first calorimeter plane);
- T & P1 & P2 (trigger plane and first two calorimeter planes).

The three rate meters as a function of L-shell (McIlwain parameter [11]) and time are shown for the period from August to September 2018; in this time interval a geomagnetic G3-class storm occurred. The initial phase of the storm started on late Aug 25th (as it can be inferred

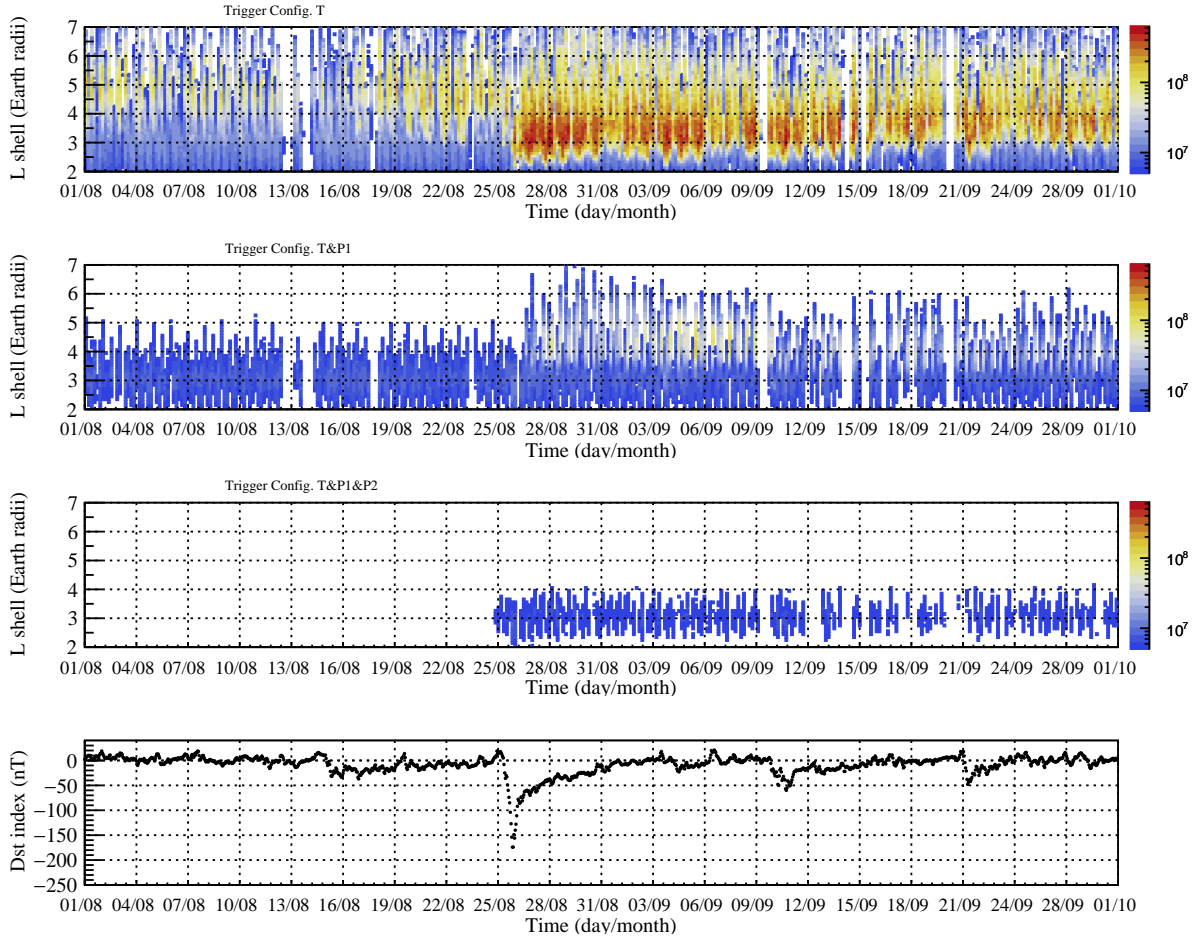


Figure 19: Rate meter of three different trigger configurations during the G3 geomagnetic storm of August 25-26, 2018 (marked by strong decrease in the Dst index shown in the bottom panel) increasing the number of planes used for the generation of the trigger.

1136 from the trend of the Dst index in the bottom panel of fig₁₁₅₅
 1137 ure 19) and the injection of low-energy particles is visible₁₁₅₆
 1138 from the increase in counts detected by the first rate meter₁₁₅₇
 1139 (figure 19 top panel). Adding more planes in the trigger₁₁₅₈
 1140 configuration results in increasing the energy threshold for₁₁₅₉
 1141 the detection, thus reducing the particle rate; that is why₁₁₆₀
 1142 the second and third rate meters (figure 19 central panels)₁₁₆₁
 1143 only detect very faint increases at storm onset. ₁₁₆₂

1144 So far, HEPD has been functioning in orbit as expected₁₁₆₃
 1145 after 33 months of lifetime. The measurement of the galac-₁₁₆₄
 1146 tic protons flux in the interval 40–250 MeV has been suc-₁₁₆₅
 1147 cessfully concluded, as reported in [12]. ₁₁₆₆

1148 7. Conclusion

1149 In this article, we have described the electronics archi-₁₁₆₉
 1150 tecture of the High Energy Particle Detector (HEPD)₁₁₇₀
 1151 an instrument that has been flying on board the Chinese₁₁₇₁
 1152 CSES satellite since February 2018. HEPD electronics has₁₁₇₂
 1153 been designed according to the general requirements im-₁₁₇₃
 1154 posed by satellite operations: in this specific case, limited₁₁₇₄

power (43 W), limited temperature operating range (op-
 erating temperature: $-10\text{ }^{\circ}\text{C} \div +35\text{ }^{\circ}\text{C}$), limits on max
 number of data transferable to the ground per day (data
 budget: 50 Gb/day), and scheduled satellite lifetime of 5
 years.

Both housekeeping and scientific data have allowed us to
 verify that the instrument is working according to nominal
 in-flight parameters, and that no substantial criticality has
 been observed in the first 33 months of life. The measure-
 ment of the galactic protons flux in the interval $40 \div 250$
 MeV has been successfully concluded, and data analysis is
 continuing in order to fulfill the scientific objectives of the
 mission.

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1178 References

- 1179 [1] X. Shen, *et al.*, *The state-of-art of the China Seismo-*
1180 *Electromagnetic Satellite mission*, Science China Technological
1181 Sciencies, vol. 61, no. 5, pp.634-642, May 2018.
- 1182 [2] S. Alexandrin, *et al.*, *High-energy charged particle bursts in the*
1183 *near-Earth space as earthquake precursors*, Ann. Geophys. 21
1184 (2003) 597.
- 1185 [3] V. Sgrigna, *et al.*, *Correlations between earthquakes and*
1186 *anomalous particle bursts from SAMPEX/PET satellite obser-*
1187 *vations*, J. Atmos. Sol.-Terr. Phys. 67 (15) (2005) 1448.
- 1188 [4] R. Battiston, *et al.*, *First evidence for correlations between*
1189 *electron fluxes measured by NOAA-POES satellites and large*
1190 *seismic events*, Nuclear Phys. B Proc. Suppl. 244 (2011) 249.
- 1191 [5] P. Picozza, *et al.*, *Scientific Goals and In-orbit Performance*
1192 *of the High-energy Particle Detector on Board the CSES*, The
1193 *Astrophysical Journal Supplement Series*, vol. 243, pp. 16, July
1194 2019.
- 1195 [6] G. Ambrosi, *et al.*, *The HEPD particle detector of the CSES*
1196 *satellite mission for investigating seismo-associated perturba-*
1197 *tions of the Van Allen belts*, Sci China Tech Sci, 2018, 61:
1198 643-652.
- 1199 [7] A. Sotgiu, *et al.*, *Control and Data Acquisition Software of*
1200 *the High-Energy Particle Detector on board the CSES Space*
1201 *Mission*, Software: Practice and Experience, in press.
- 1202 [8] V. Scotti and G. Osteria, *The electronics of the*
1203 *HEPD of the CSES experiment*, Nuclear and
1204 *Particle Physics Proceedings* 291-293 (2017) 118.
1205 <https://doi.org/10.1016/j.nuclphysbps.2017.06.024>.
- 1206 [9] Callier, S. et al., *EASIROC, an Easy & Versa-*
1207 *tile ReadOut Device for SiPM*, Physics Procedia,
1208 Volume 37, 2012, Pages 1569-1576, ISSN 1875-3892,
1209 <https://doi.org/10.1016/j.phpro.2012.02.486>.
- 1210 [10] G. Ambrosi, *et al.*, *Beam test calibrations of the HEPD detector*
1211 *on board the China Seismo-Electromagnetic Satellite*, Nuclear
1212 *Inst. and Methods in Physics Research*, A 974 (2020) 164170,
1213 September 2020.
- 1214 [11] McIlwain, *et al.*, *Coordinates for mapping the distribution of*
1215 *magnetically trapped particles*, J. Geophys. 349 Res., 66 (11),
1216 3681– 3691, 1961.
- 1217 [12] S. Bartocci, *et al.*, *Galactic Cosmic-Ray Hydrogen Spectra in*
1218 *the 40–250 MeV Range Measured by the High-energy Partic-*
1219 *le Detector (HEPD) on board the CSES-01 Satellite between*
1220 *2018 and 2020*, The Astrophysical Journal, 901:8 (7pp), 2020
1221 September 20.